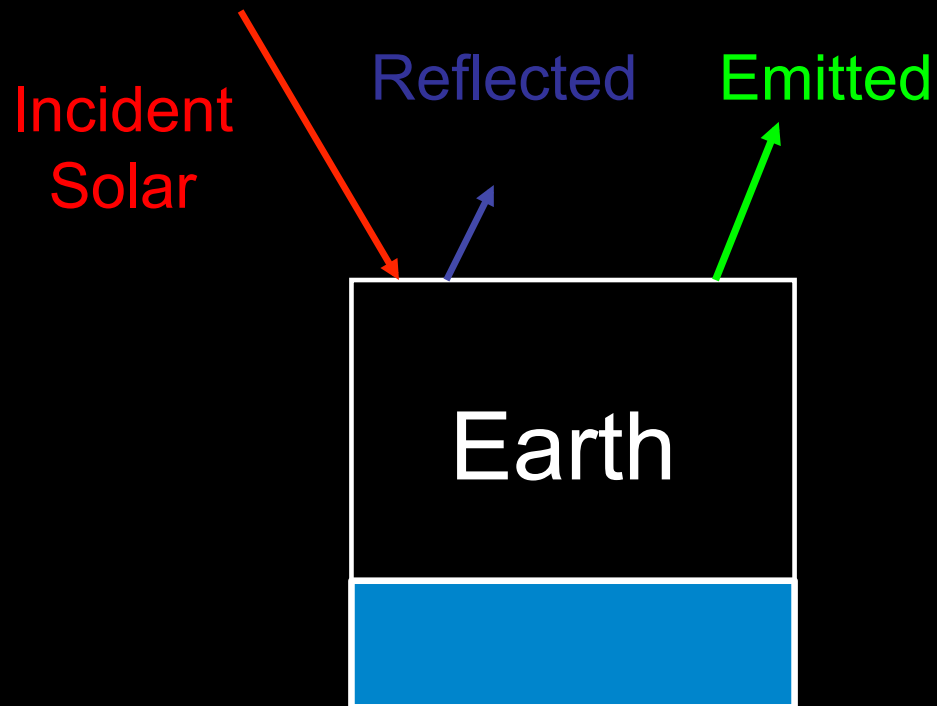


# Global Scale Energy Fluxes: Comparison of Observational Estimates and Model Simulations

Aaron Donohoe -- MIT  
David Battisti -- UW  
CERES Science Team Meeting  
4/23/2014

# Global mean energy balance

- The Earth receives energy from the sun (and reflects back some portion of it)
- To come into energy balance (equilibrium) the Earth must emit the same amount of energy it receives

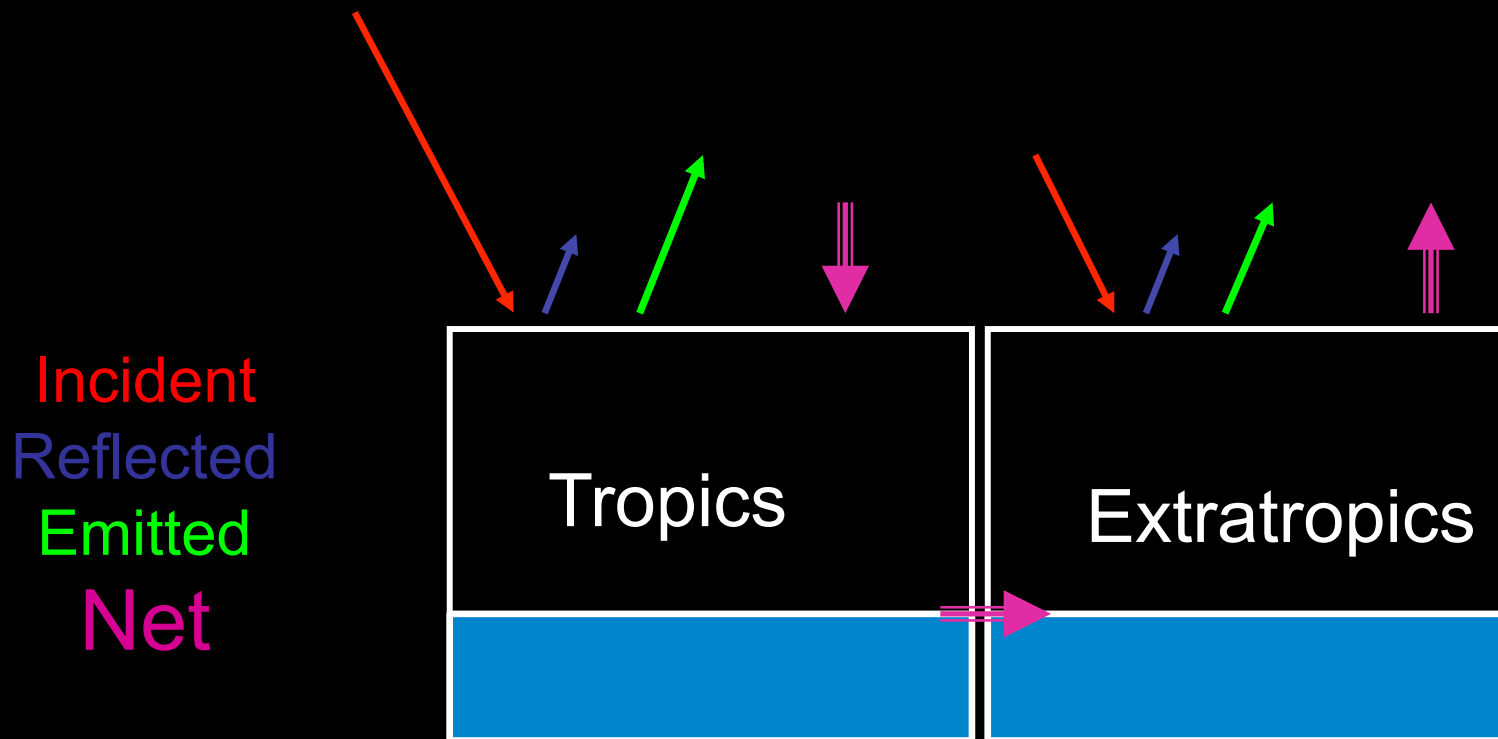


$$\text{Incident} - \text{Reflected} = \text{Emitted}$$

$$S(1-\alpha_p) = \sigma T_e^4$$

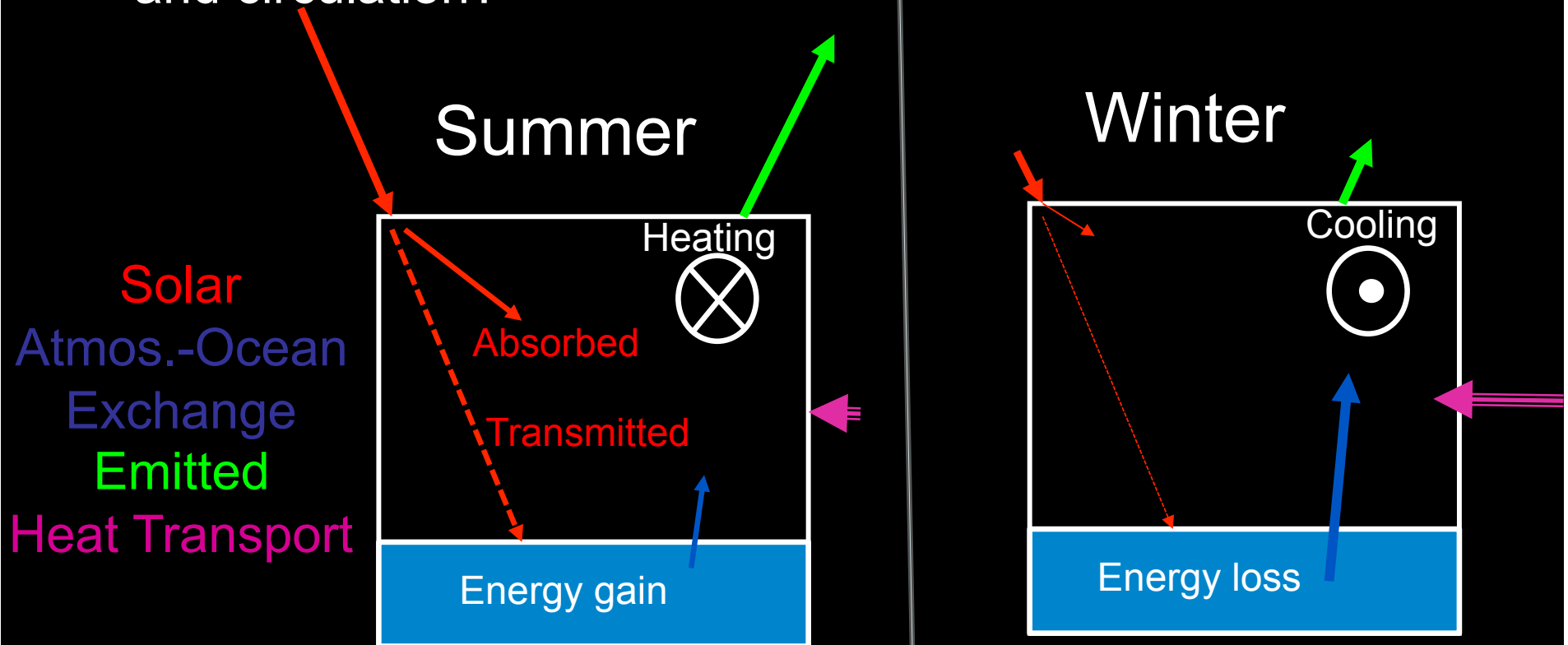
# Equator-to-pole contrast

- The tropics receive more solar radiation than the high latitudes (extratropics)
- To come to equilibrium, the tropics must either emit excess radiation or transport energy to the extratropics



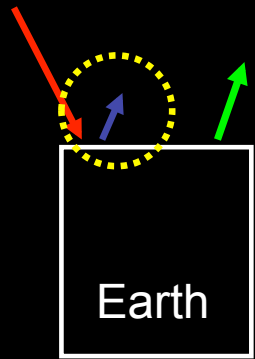
# Seasonal Cycle

- Seasonal variations in solar insolation are of order 1 in the extratropics
- How much energy goes into the atmosphere versus the ocean to drive seasonal variations in temperature and circulation?



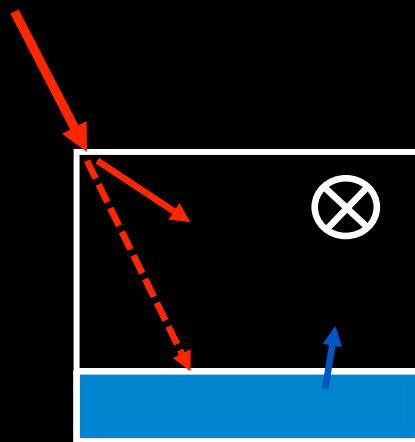
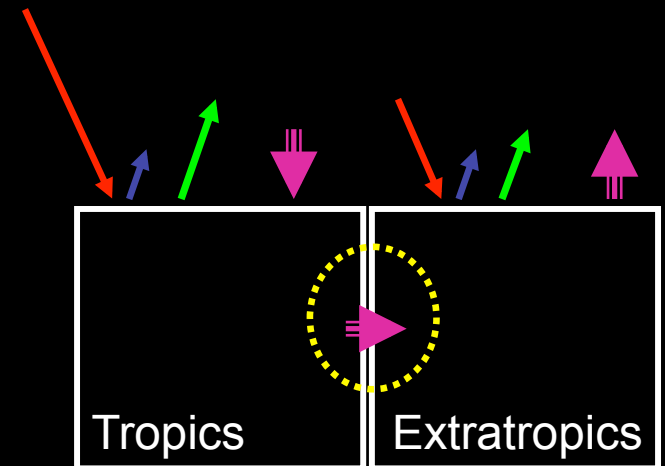


# Outline



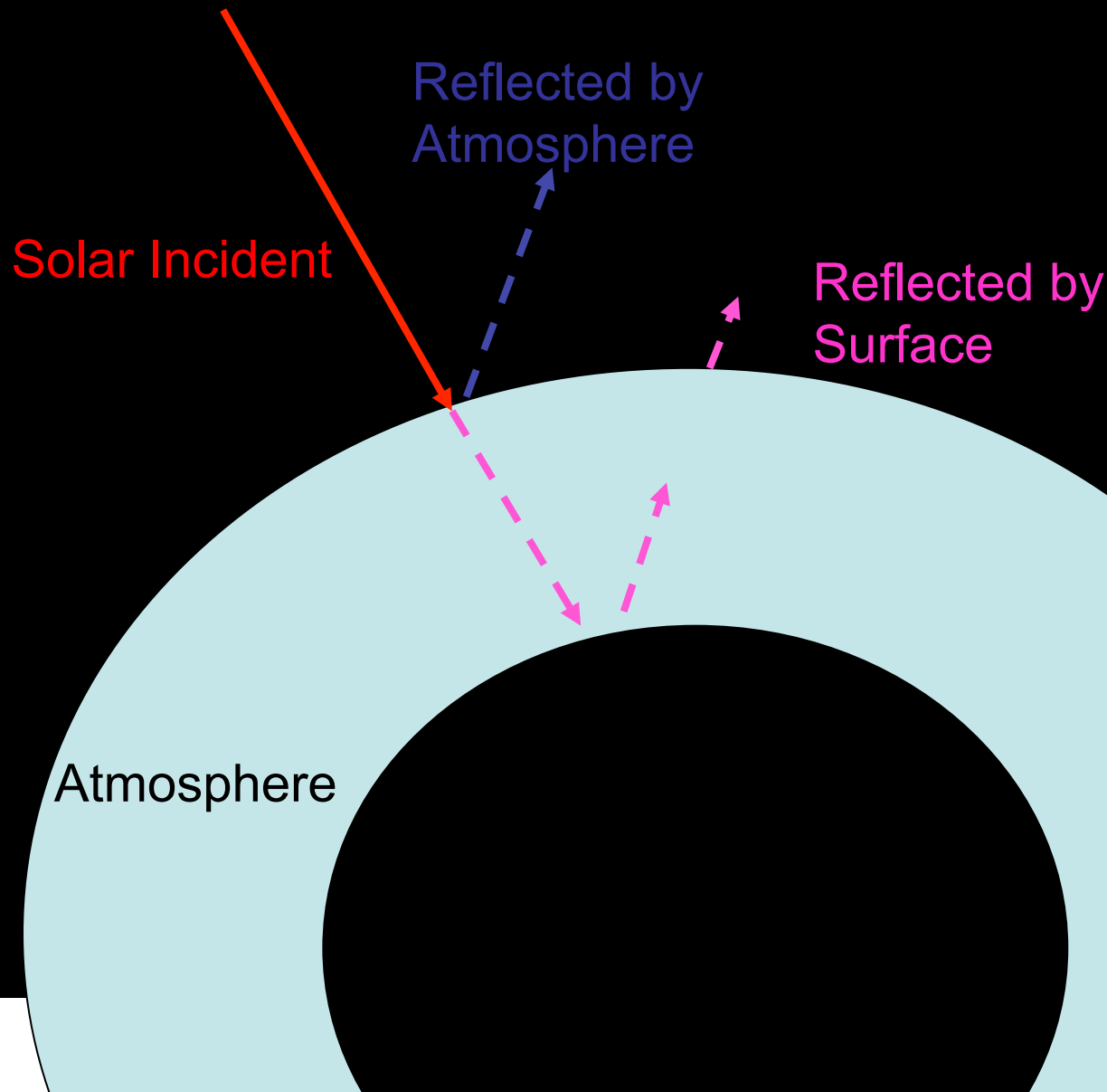
1. What determines the Earth's planetary albedo? (How much solar radiation gets reflected)

2. What determines the meridional heat transport in the climate system?



3. How do seasonal variations in solar insolation lead to atmospheric heating?

# 1 : What determines the Earth's planetary albedo? (solar radiation reflected at top of atmosphere)



# Simplified (isotropic) shortwave radiation model

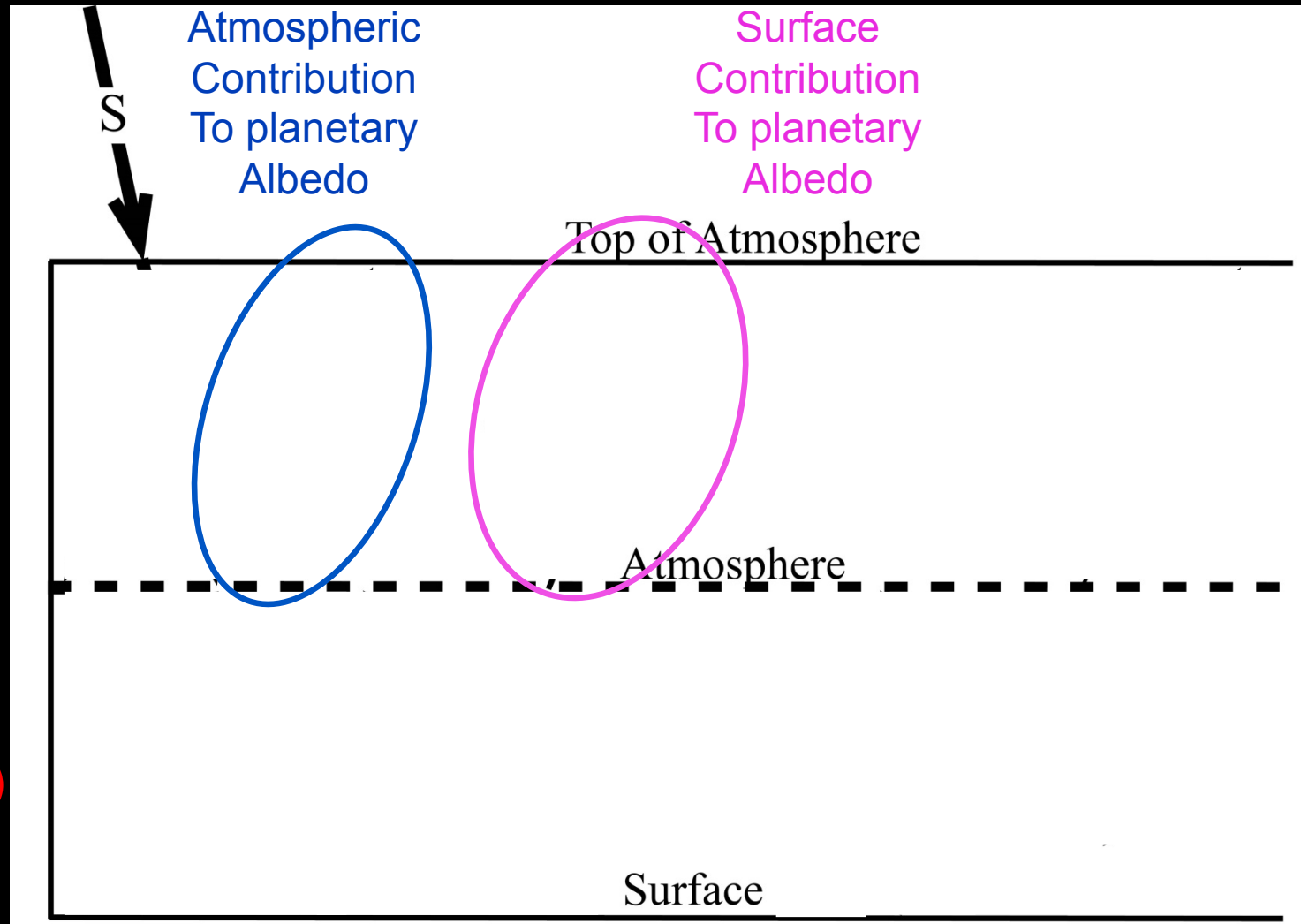
$S$  = incident

$R$  = cloud reflection

$A$  = absorption

$\alpha$  = surface albedo

(UNKNOWN)

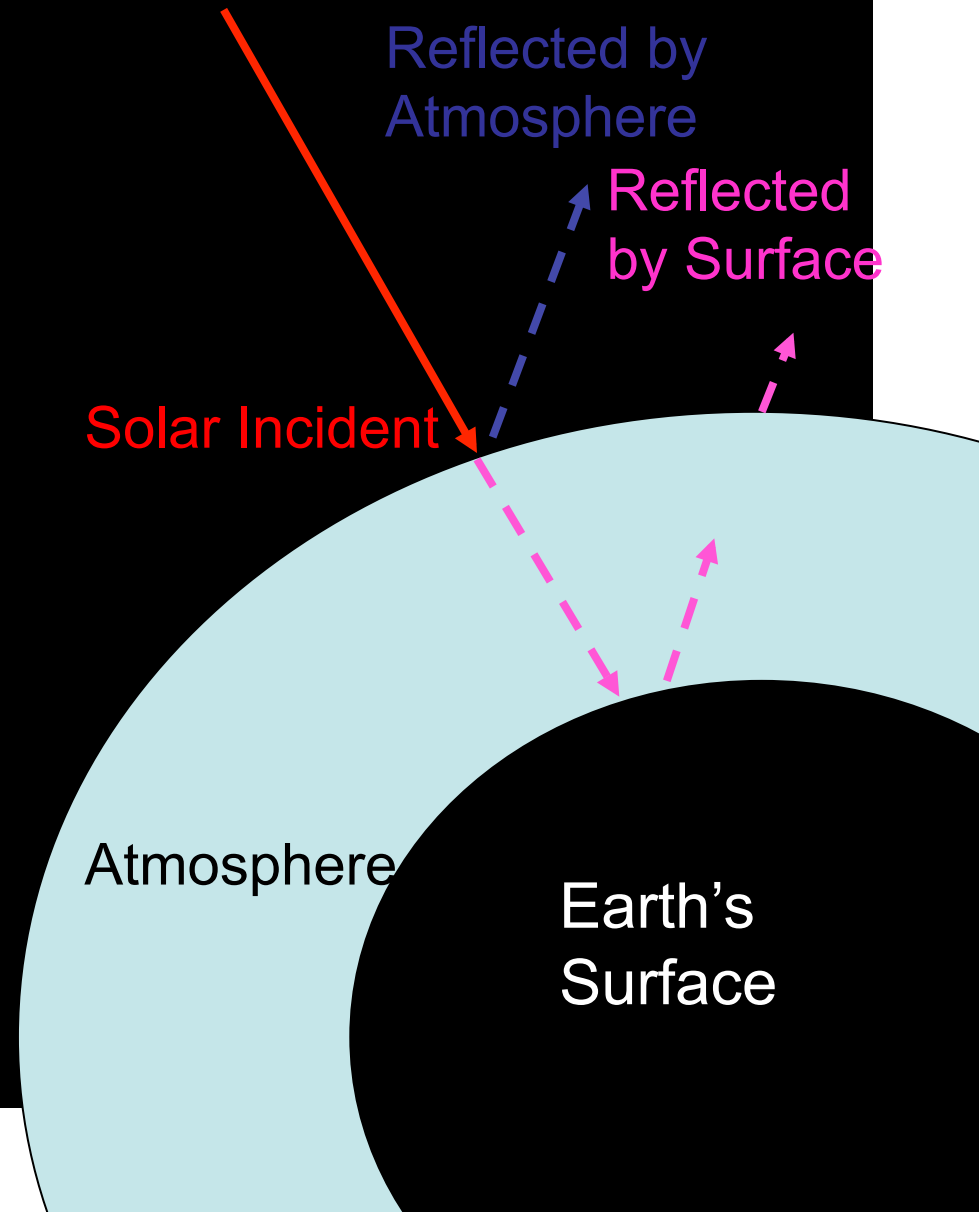


# Partitioning of planetary albedo into atmospheric and surface components

$$\alpha_P = \alpha_{P,ATMOS} + \alpha_{P,SURF}$$

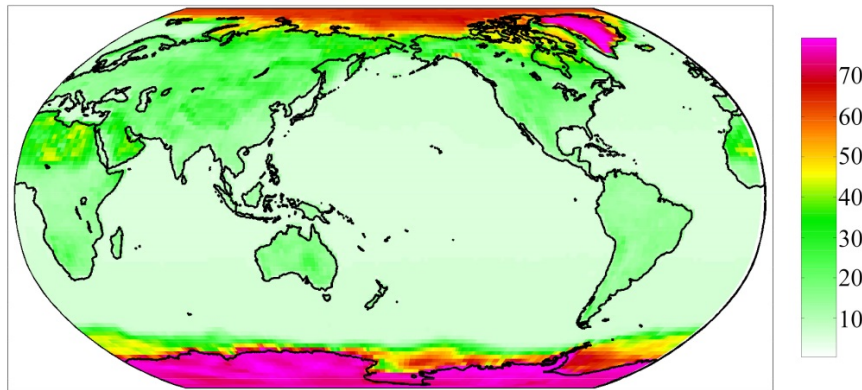
$$\alpha_{P,ATMOS} = R$$

$$\alpha_{P,SURF} = \frac{\alpha(1-R-A)^2}{(1-\alpha R)}$$

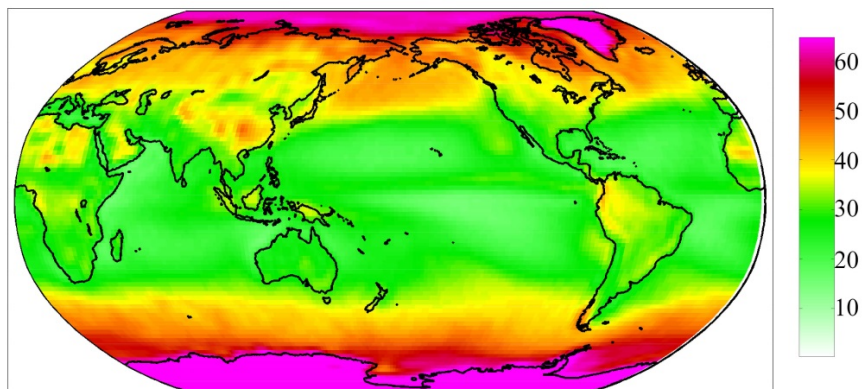


# Observed (CERES) surface and atmospheric contribution to planetary albedo

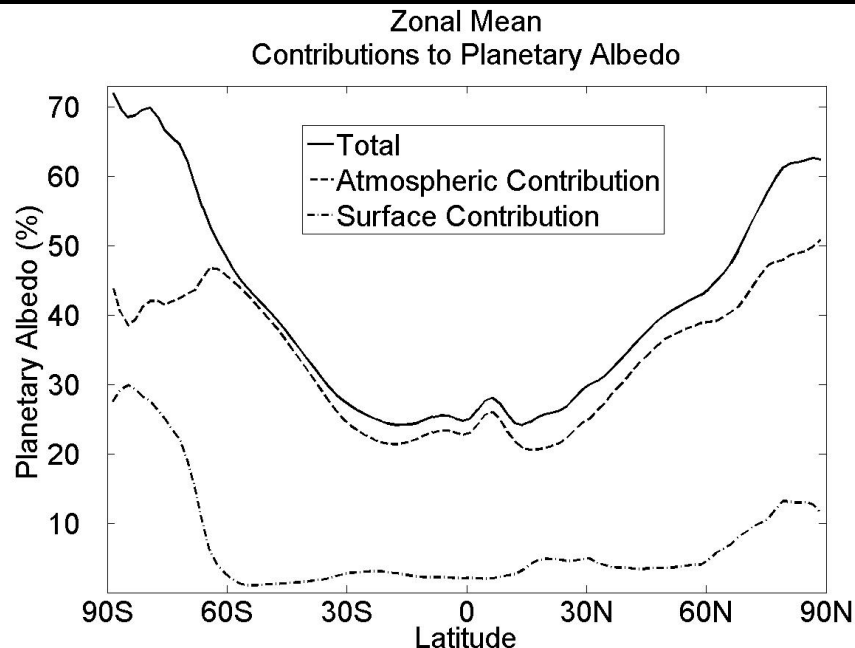
Surface Albedo (%)



Planetary Albedo (%)



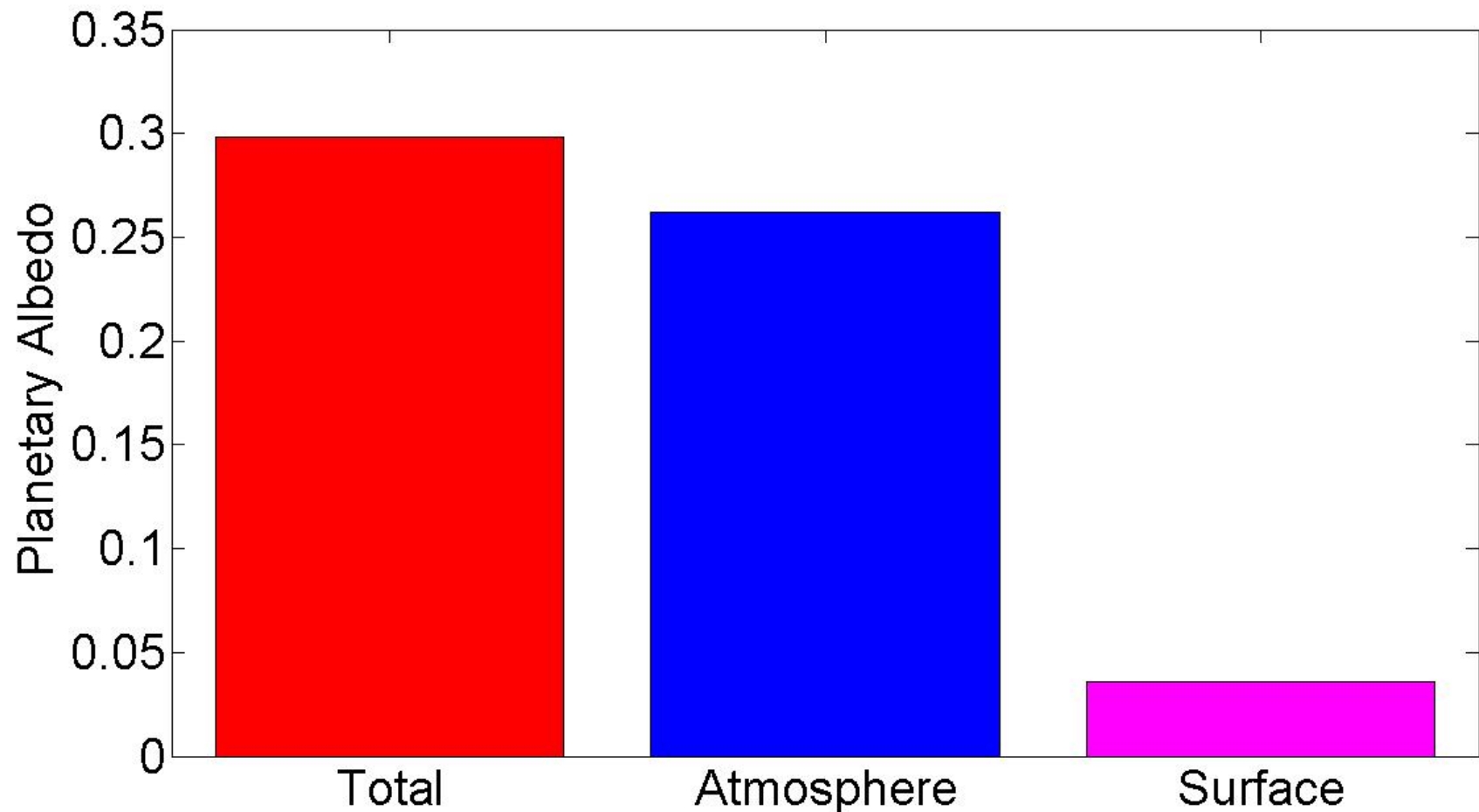
# Observed Surface and atmospheric contribution to planetary albedo



Surface albedo and surface contribution to planetary albedo ( $\alpha$  and  $\alpha_{P,SURF}$ )

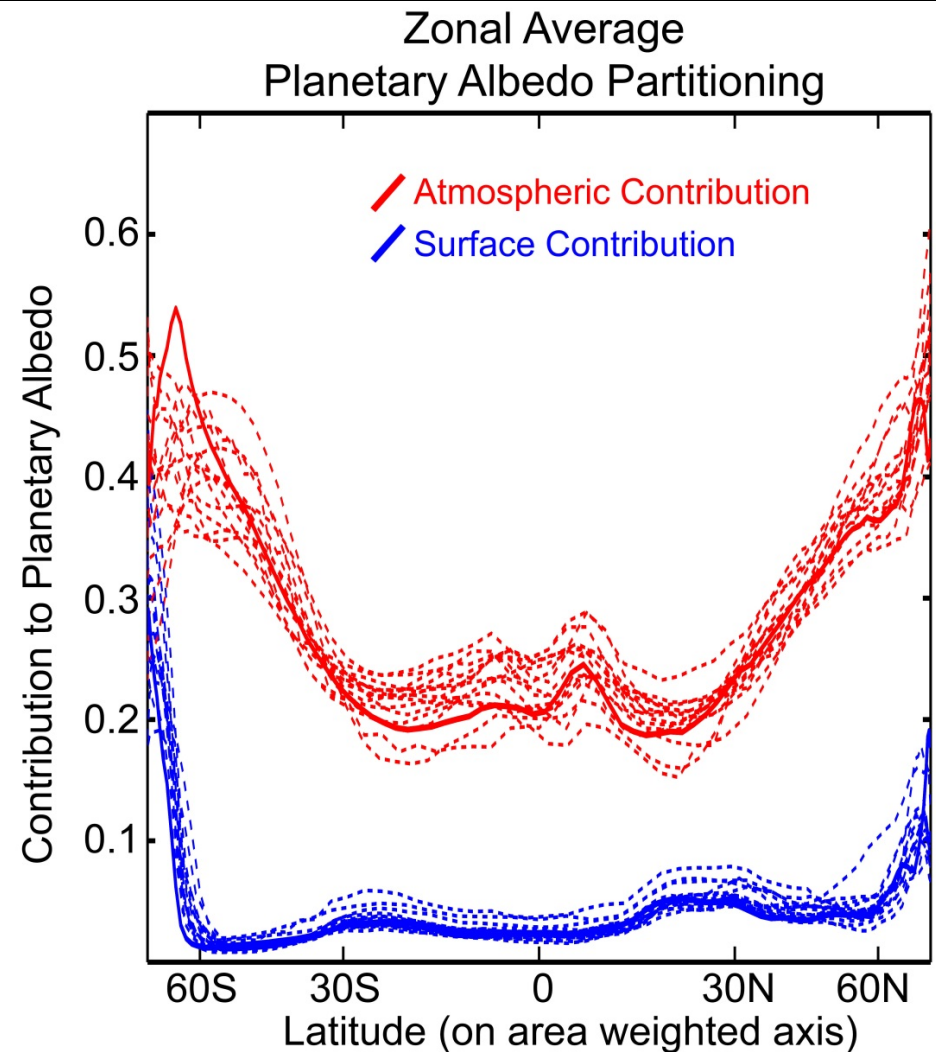
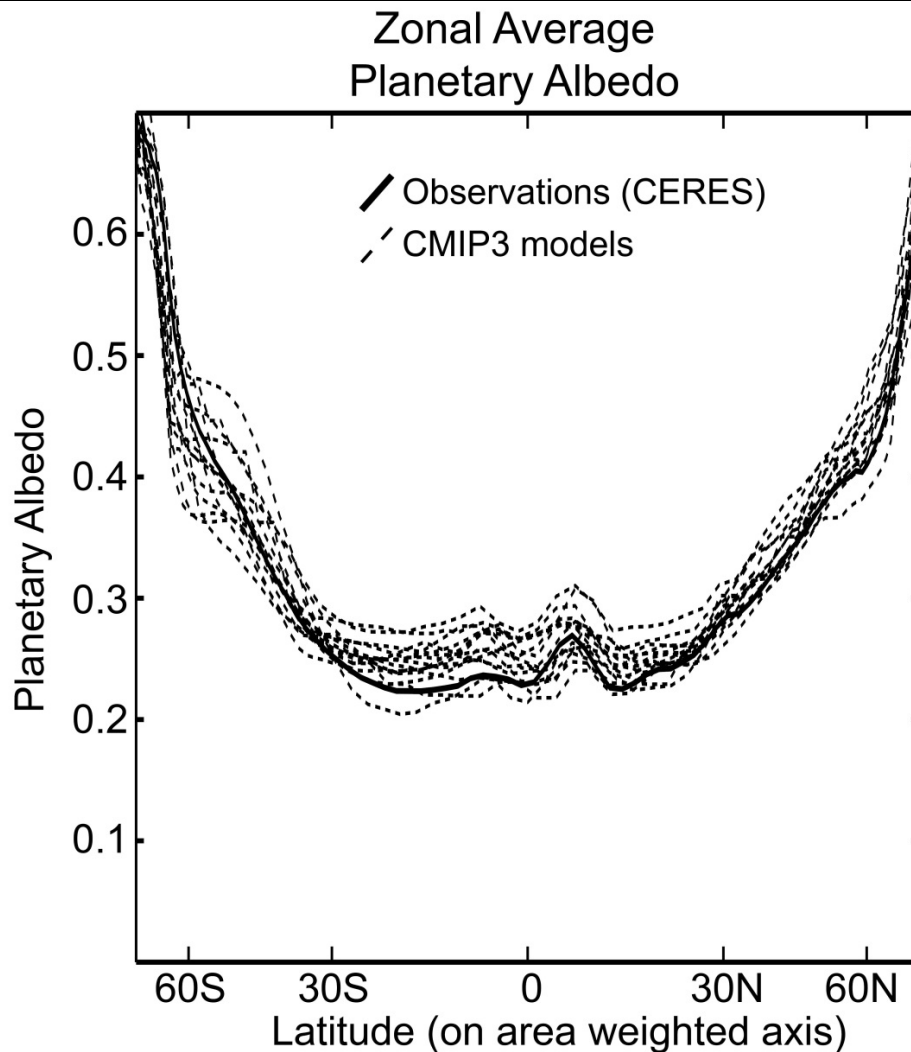
$$\alpha_{P,SURF} = \frac{\alpha(1-R-A)^2}{(1-\alpha R)}$$

# Observed Global Mean Planetary Albedo and its atmospheric/surface Partitioning

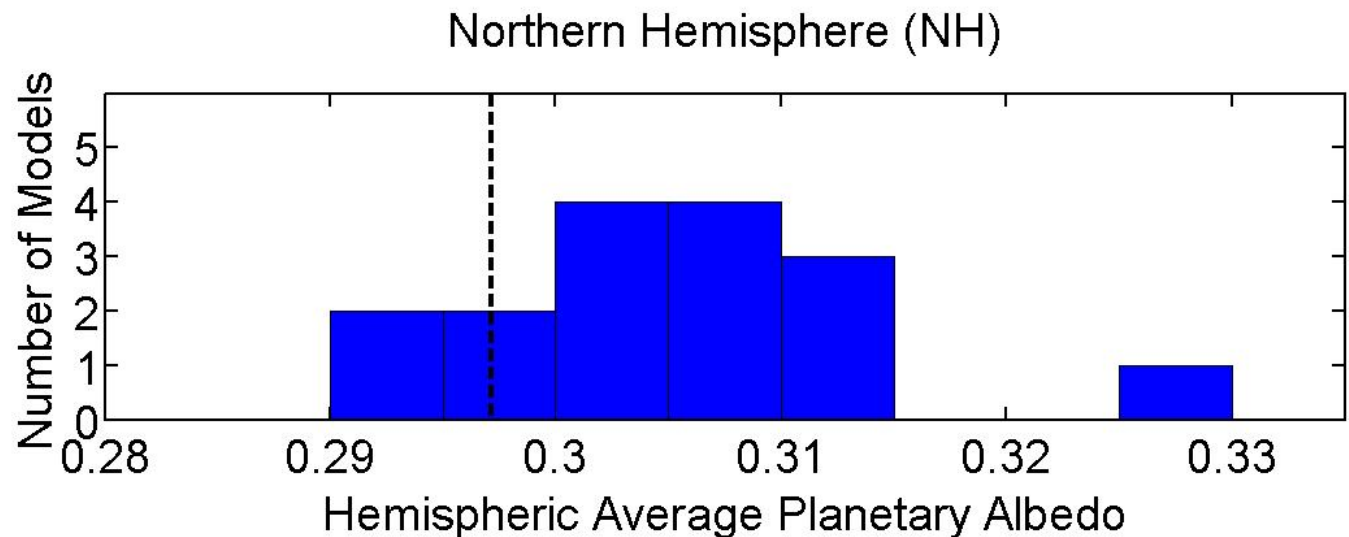
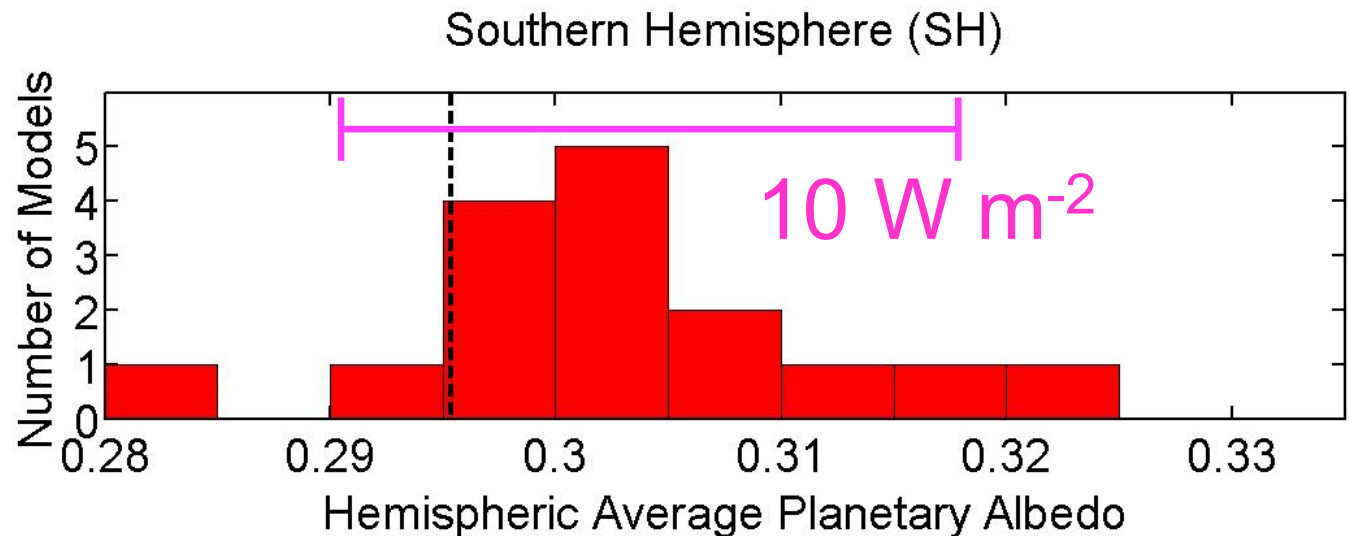




# Planetary Albedo Partitioning: Comparison of models (CMIP3 pre-industrial) and observations (CERES)

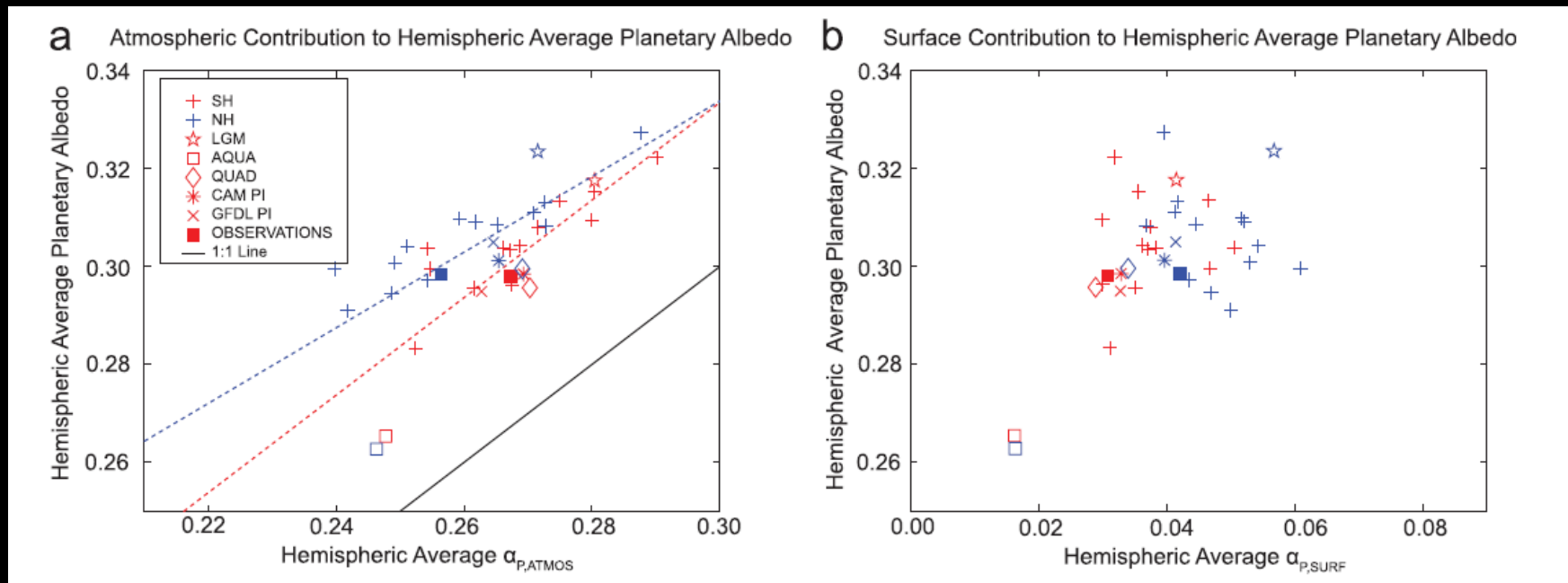


# Histogram of hemispheric average planetary albedo



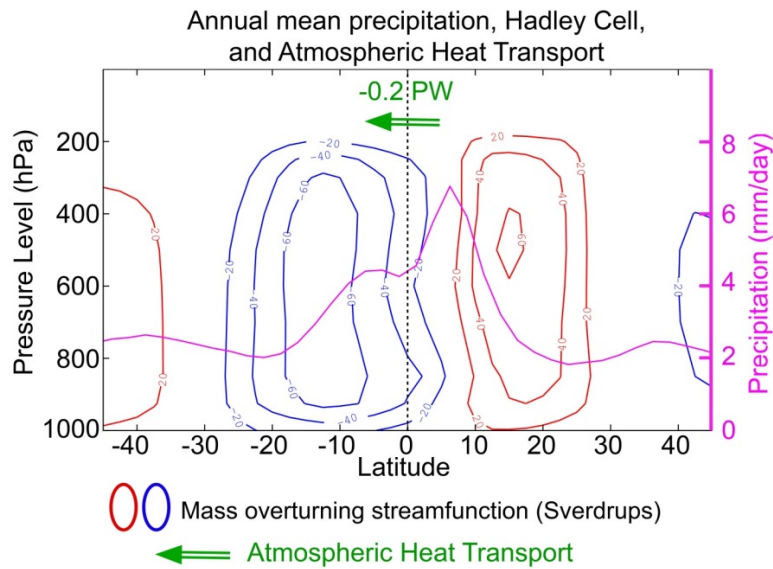
.....  
Observations  
(CERES)

# Hemispheric average planetary albedo partitioning in climate models



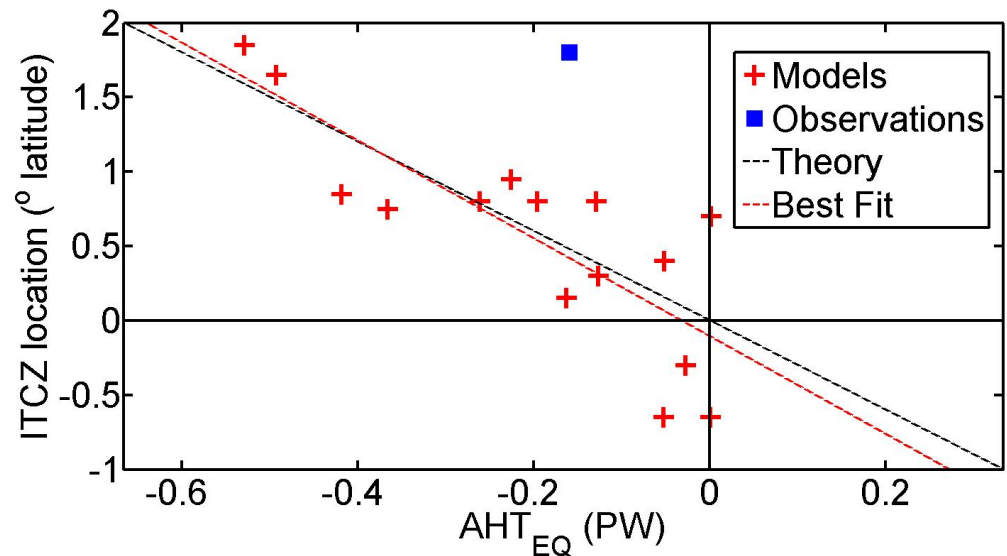
- Inter-model spread in Hemispheric average planetary albedo is a consequence of differences in cloud reflection
- Surface albedo makes a much smaller contribution to basic state albedo, model bias and inter-model spread

# Inter-hemispheric energy transport and the location of tropical precipitation

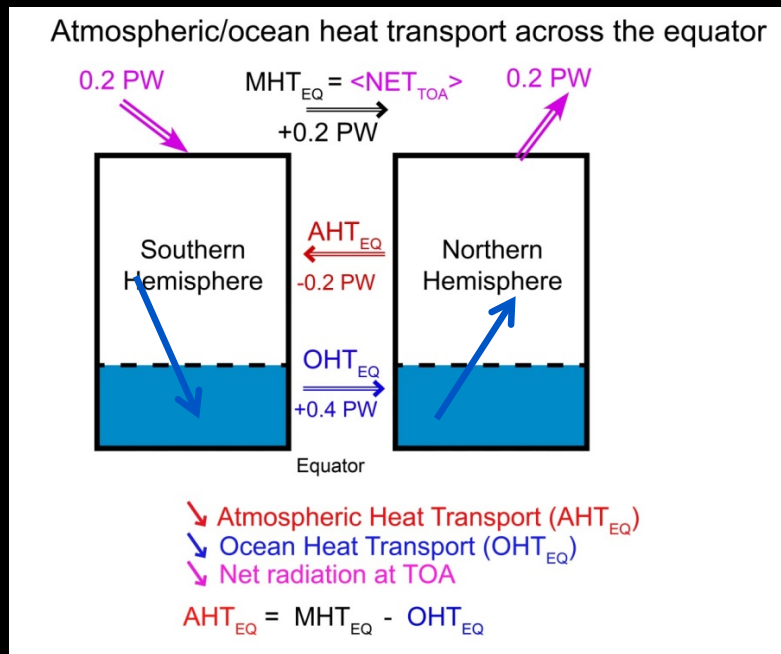


ITCZ location and atmospheric heat transport across the equator ( $AHT_{EQ}$ ) are both a consequence of the Hadley cell location  
→ The atmosphere moves energy away from the ITCZ

The inter-model spread in  $AHT_{EQ}$  and ITCZ location are strongly correlated



# Hemispheric contrast of radiation and ITCZ location

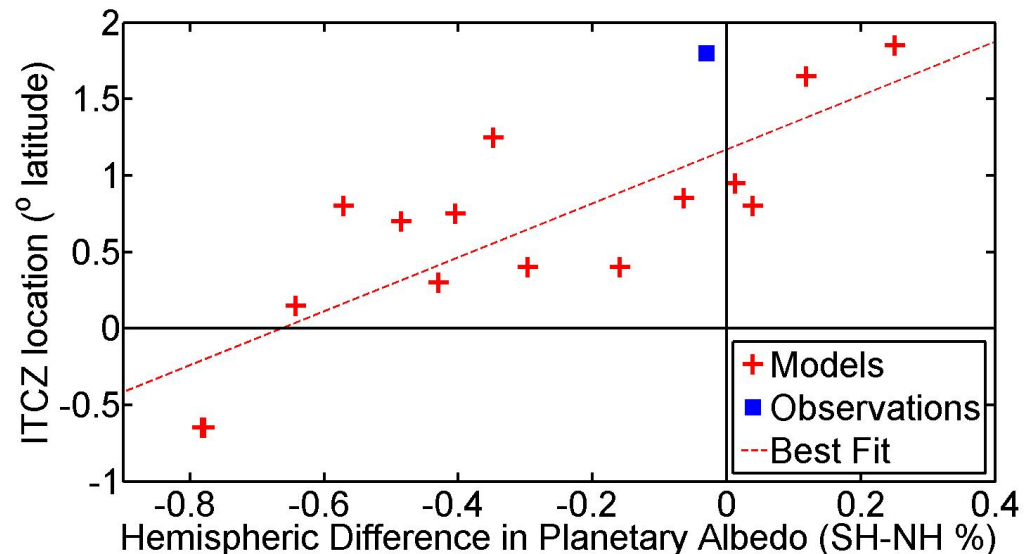


ITCZ lives in the hemisphere where the atmosphere is heated more strongly

-> In observations, the radiative input to each hemisphere is nearly equal and ITCZ is North of the equator because of Northward ocean heat transport

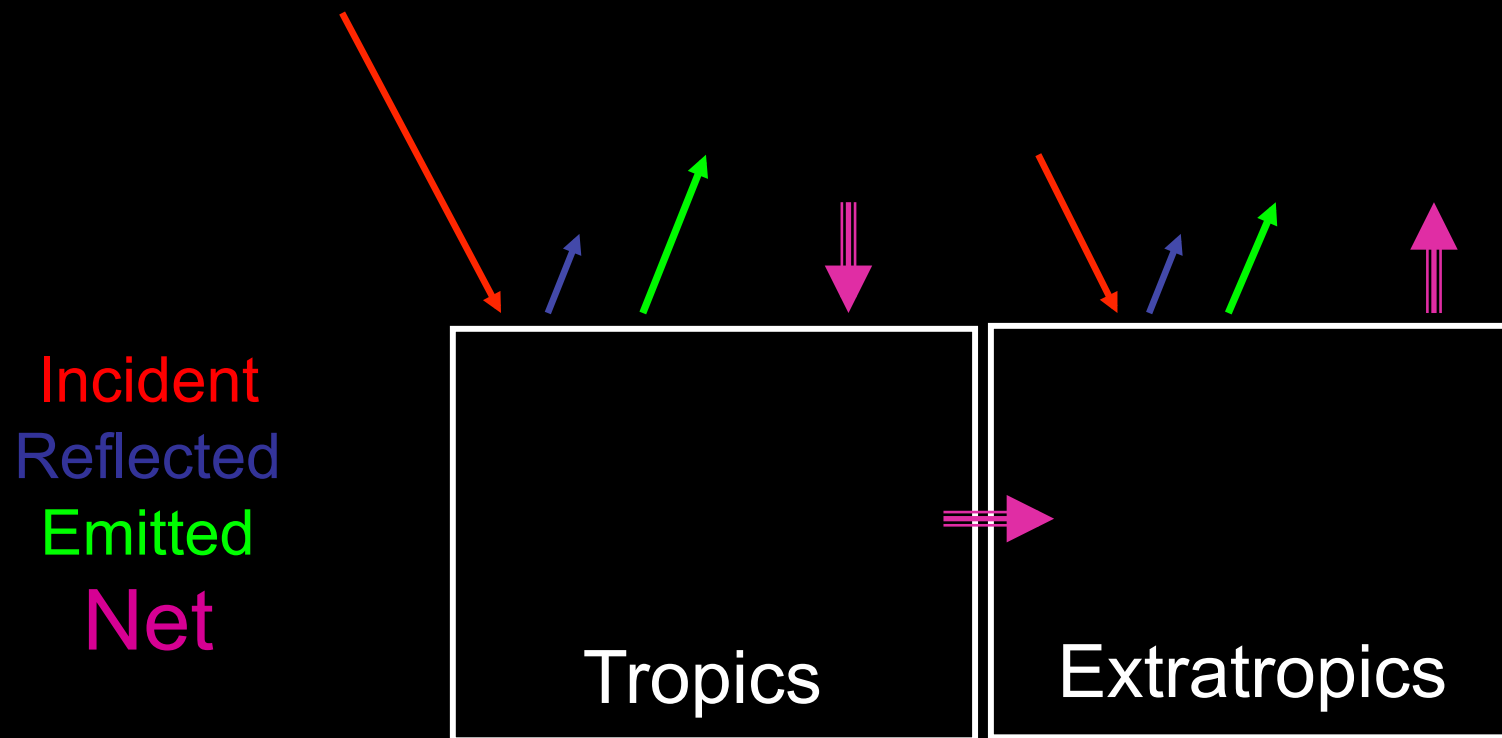
Model biases in planetary albedo lead to ITCZ biases

-> SH absorbs too much shortwave and ITCZ is too far south

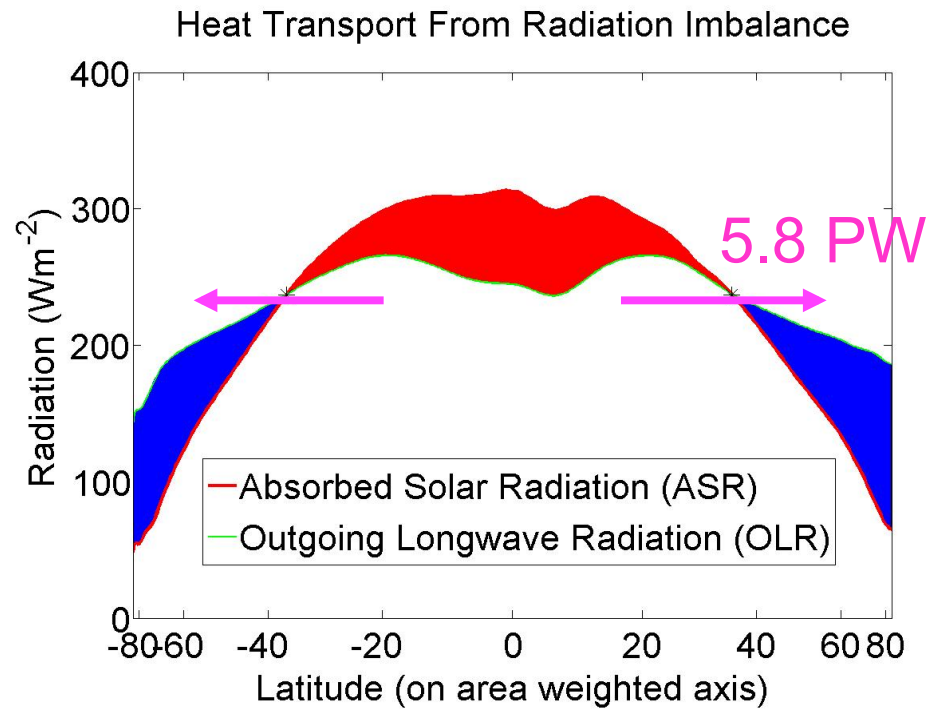


← Heating of SH

## 2 : What determines meridional heat transport?

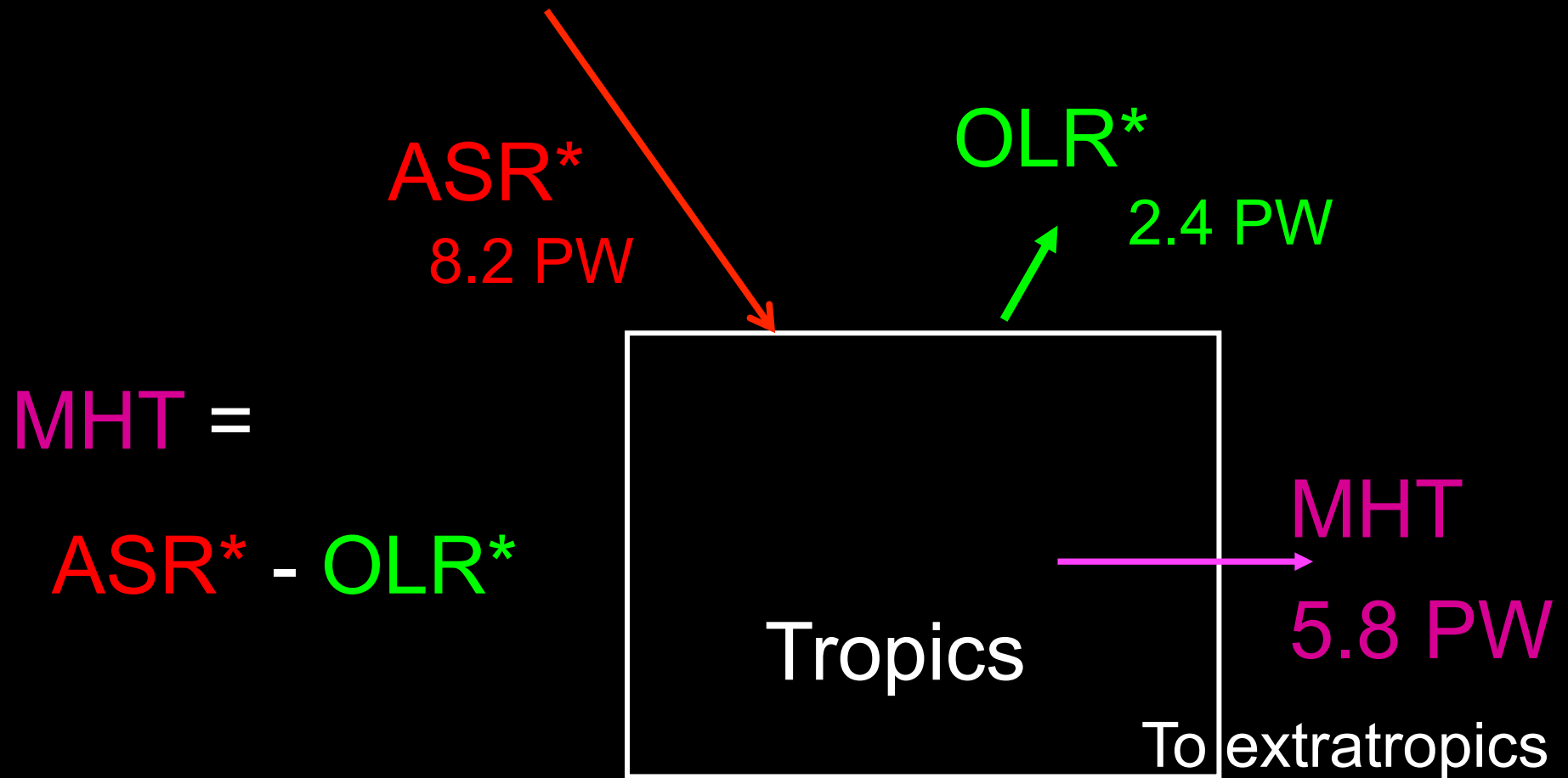


# Understanding heat transport





# ASR\*, OLR\*, MHT, and the tropical/ extratropical energy budget



All arrows are relative to the global average

ASR\*, OLR\*, MHT  
Dynamic Limit

$$\text{MHT} = \text{ASR}^*$$

$$\text{OLR}^* = 0$$

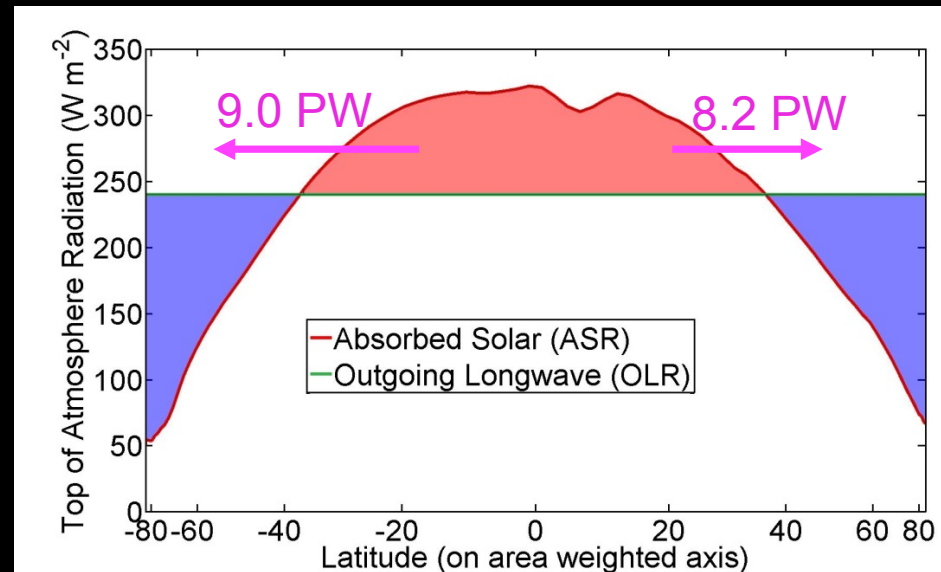
ASR\*  
8.2 PW

Tropics

MHT

8.2 PW

To extratropics

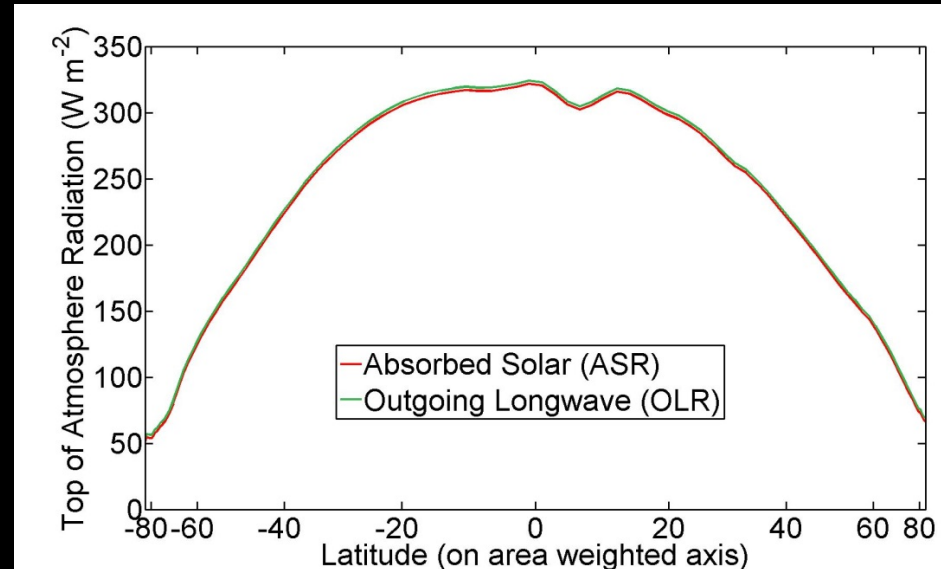
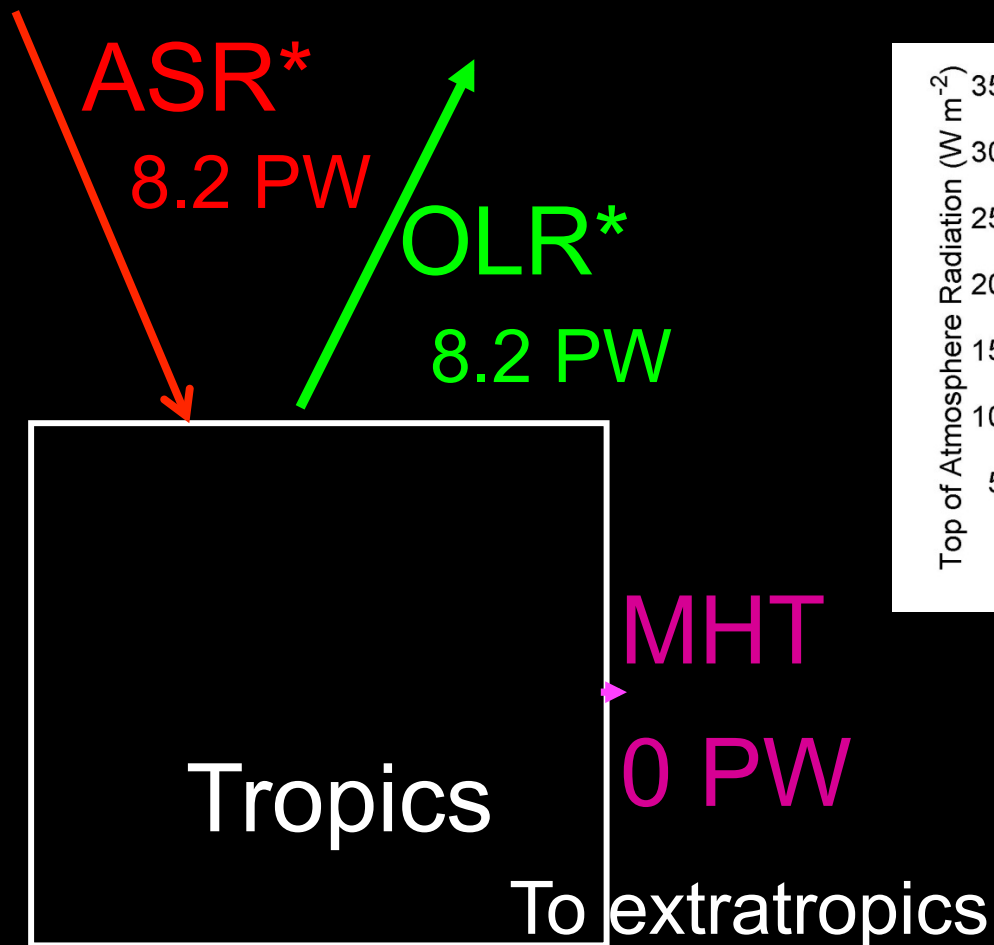


All arrows are relative to the global average

# ASR\*, OLR\*, MHT Radiative Limit

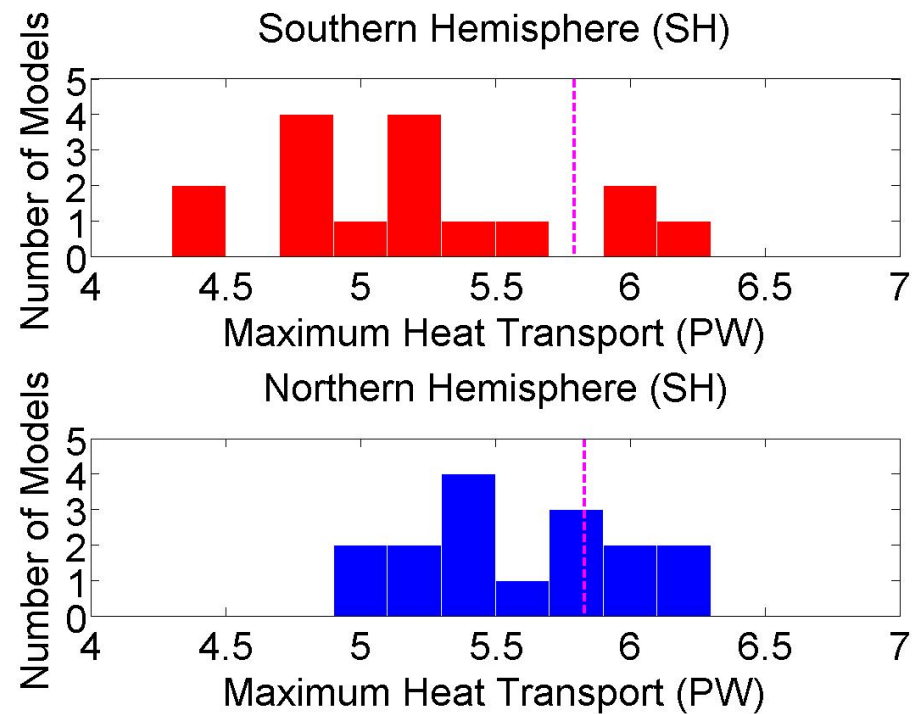
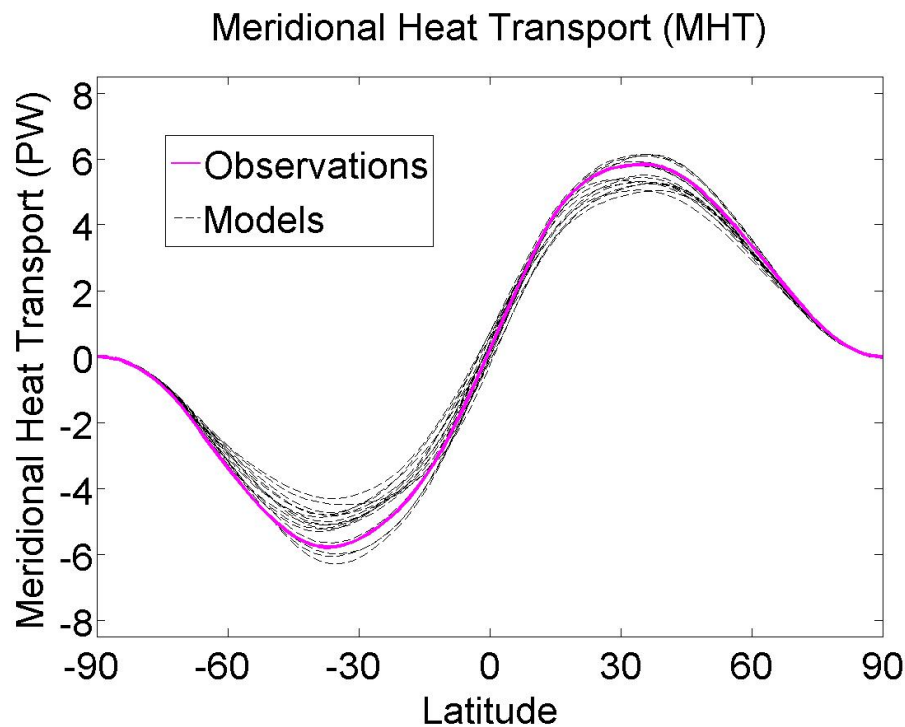
$$MHT = 0,$$

$$ASR^* = OLR^*$$



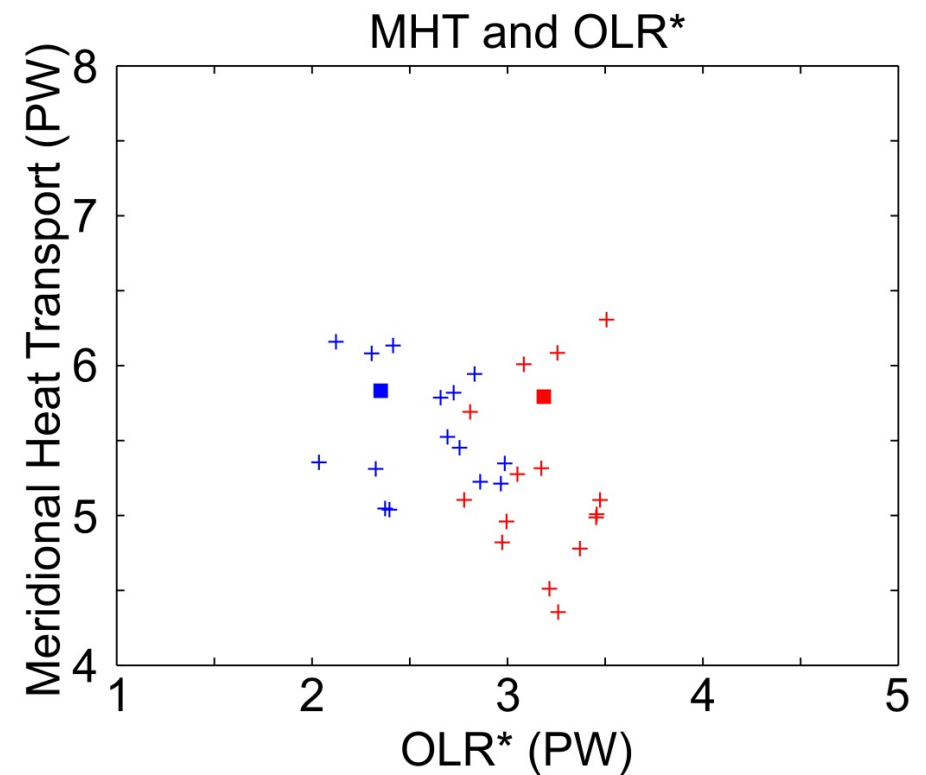
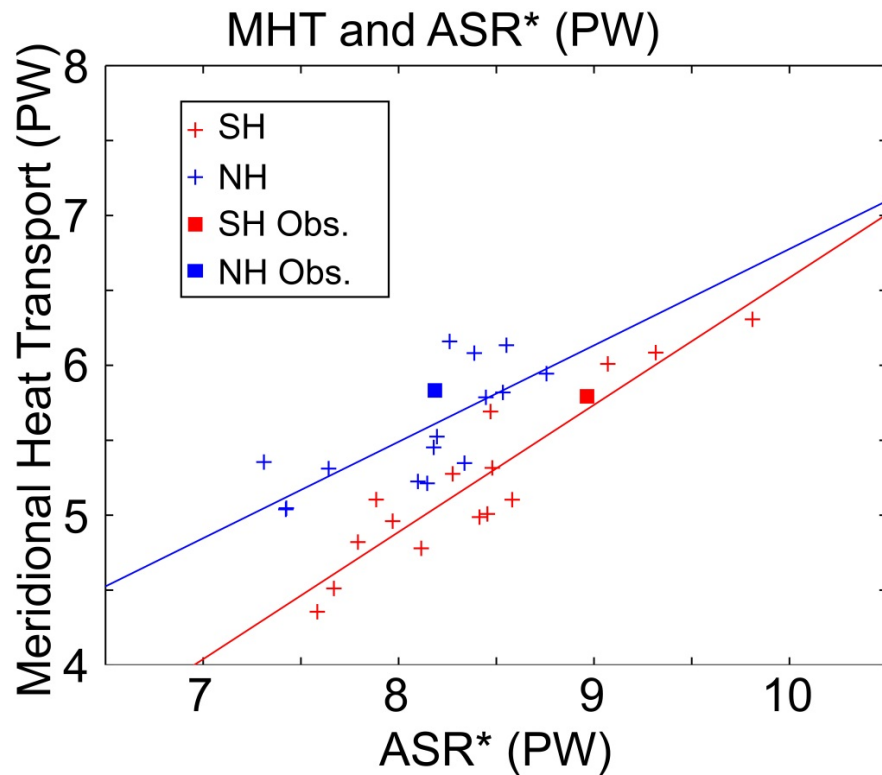
All arrows are relative to the global average

# Heat Transport In Climate Models (CMIP3)

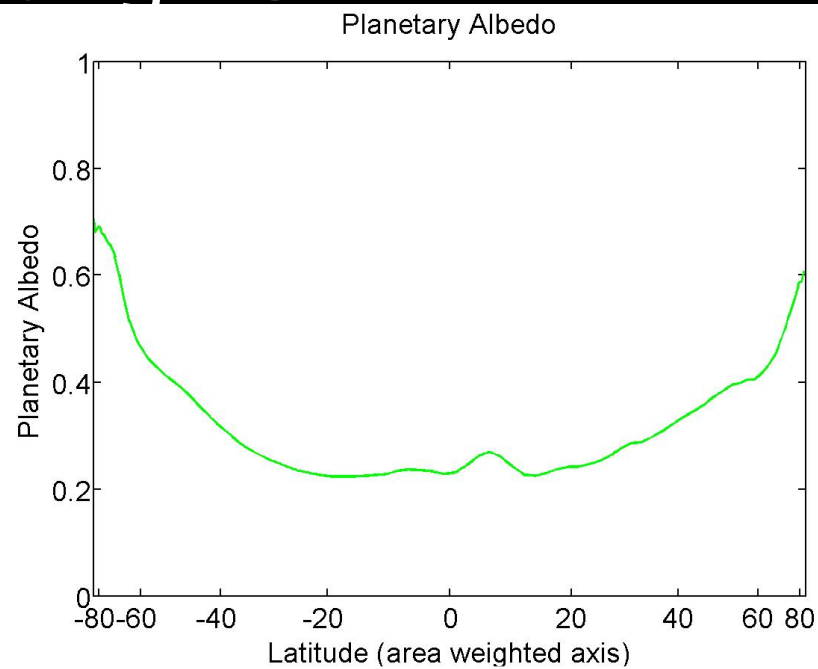
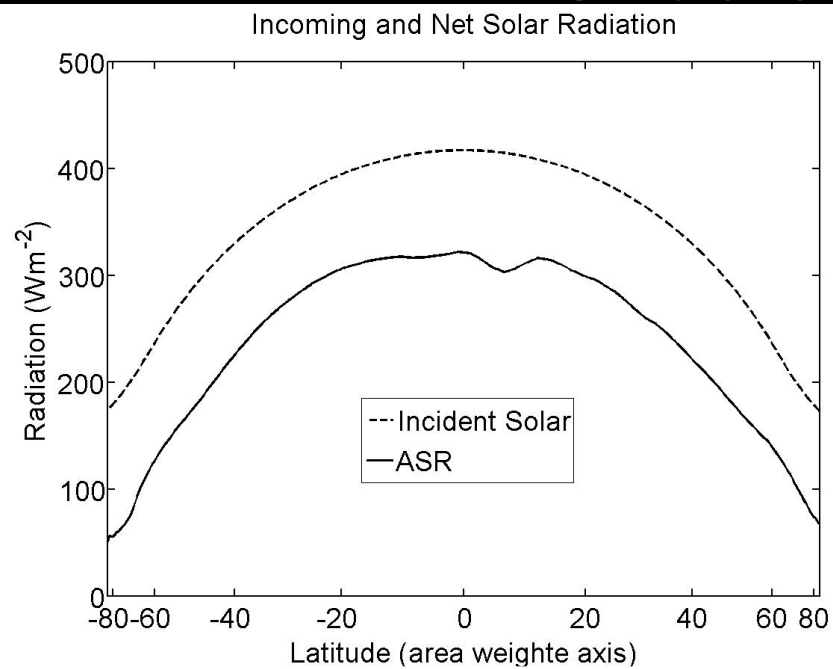


--- Observations

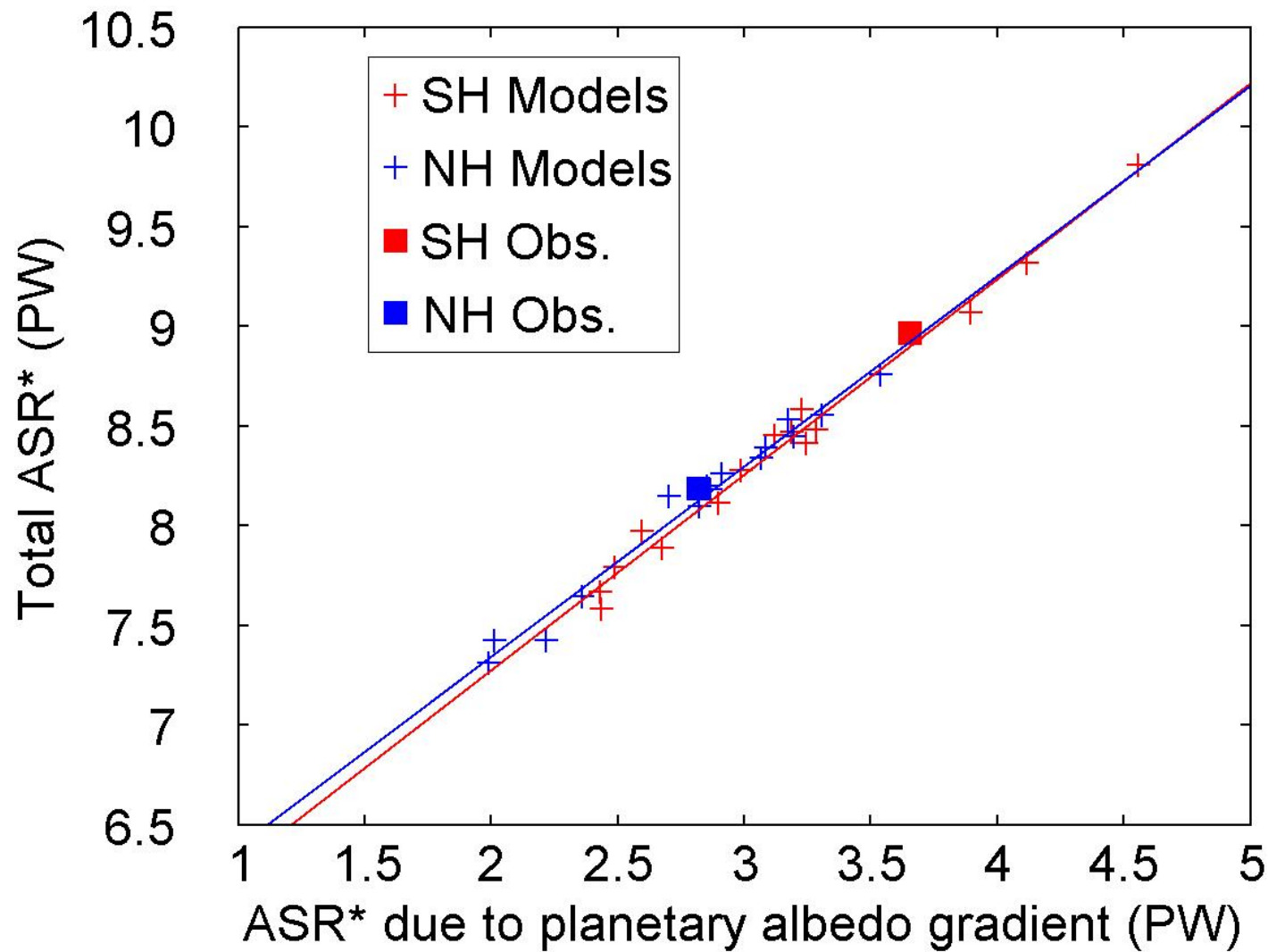
# Model heat transport spread in terms of $OLR^*$ and $ASR^*$



# Understanding ASR\*

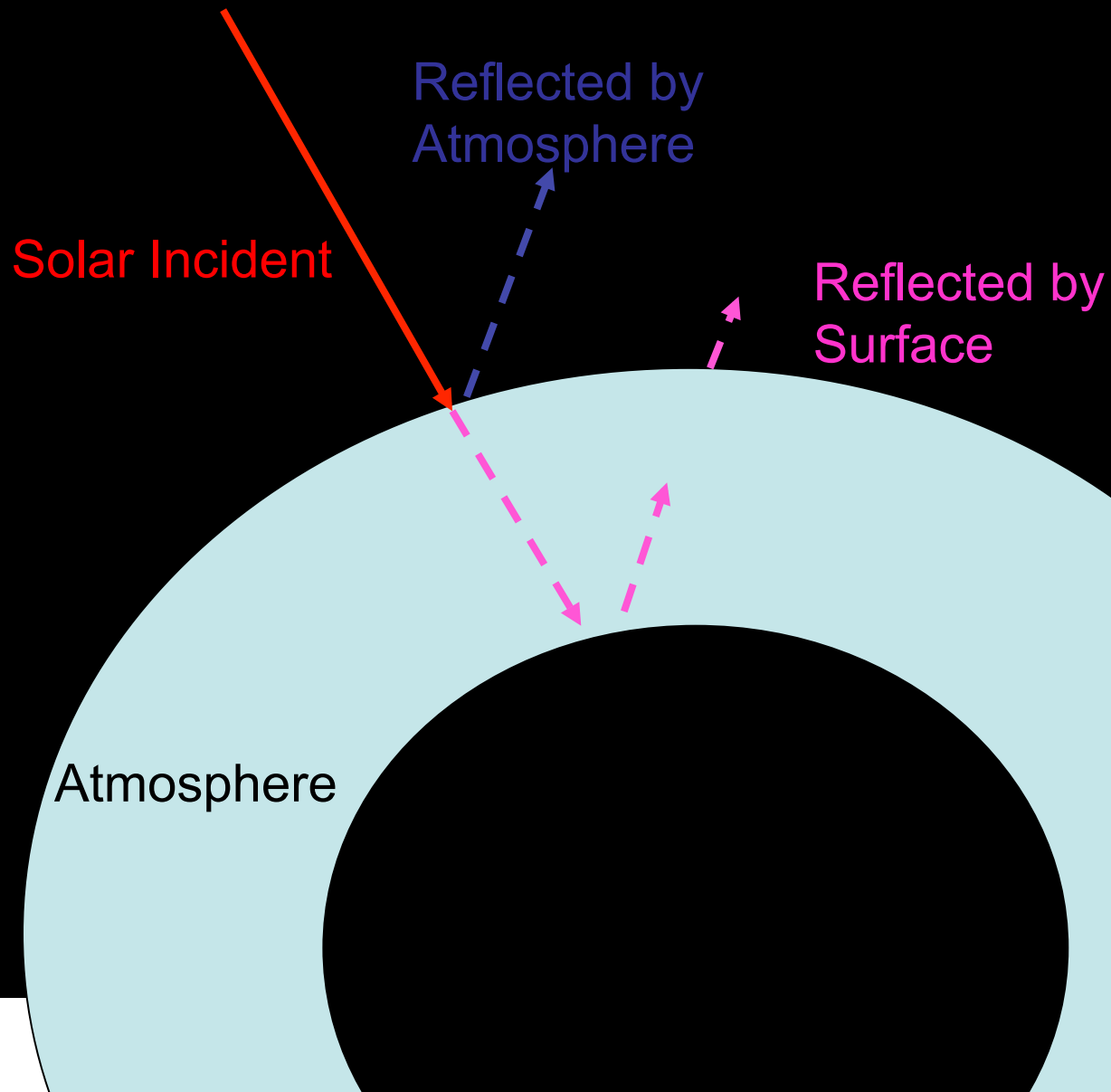


# ASR\* and planetary albedo

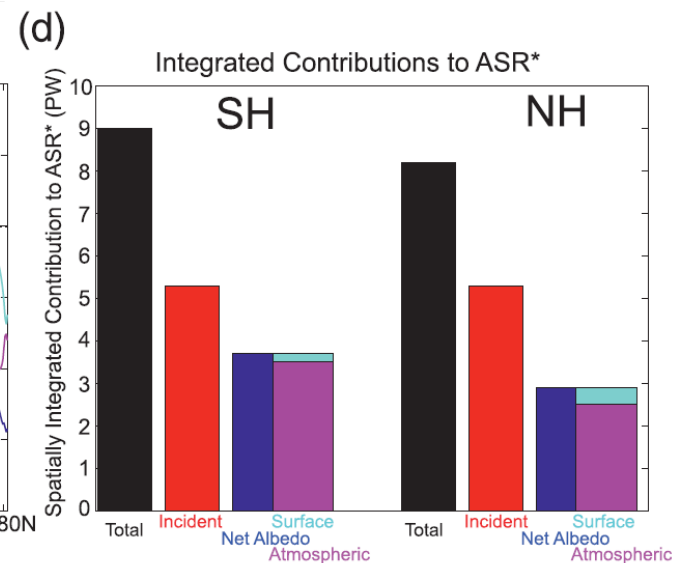
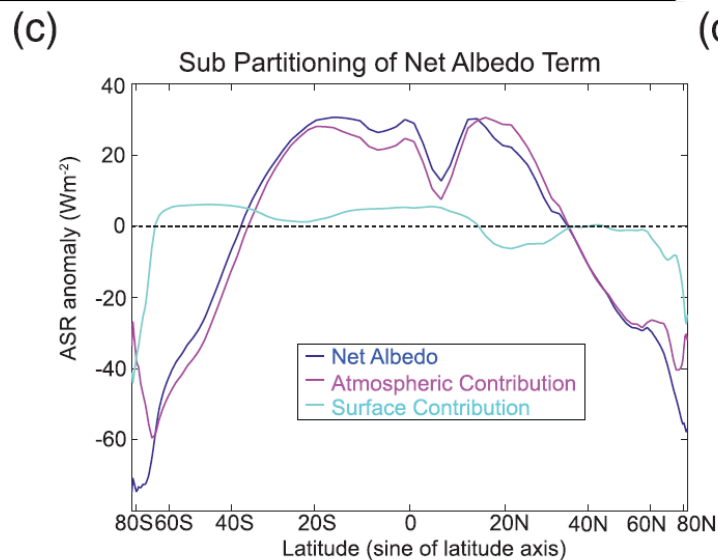
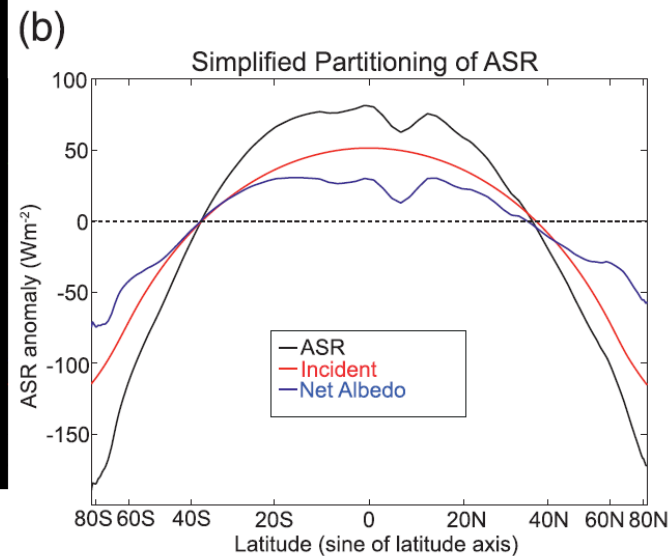




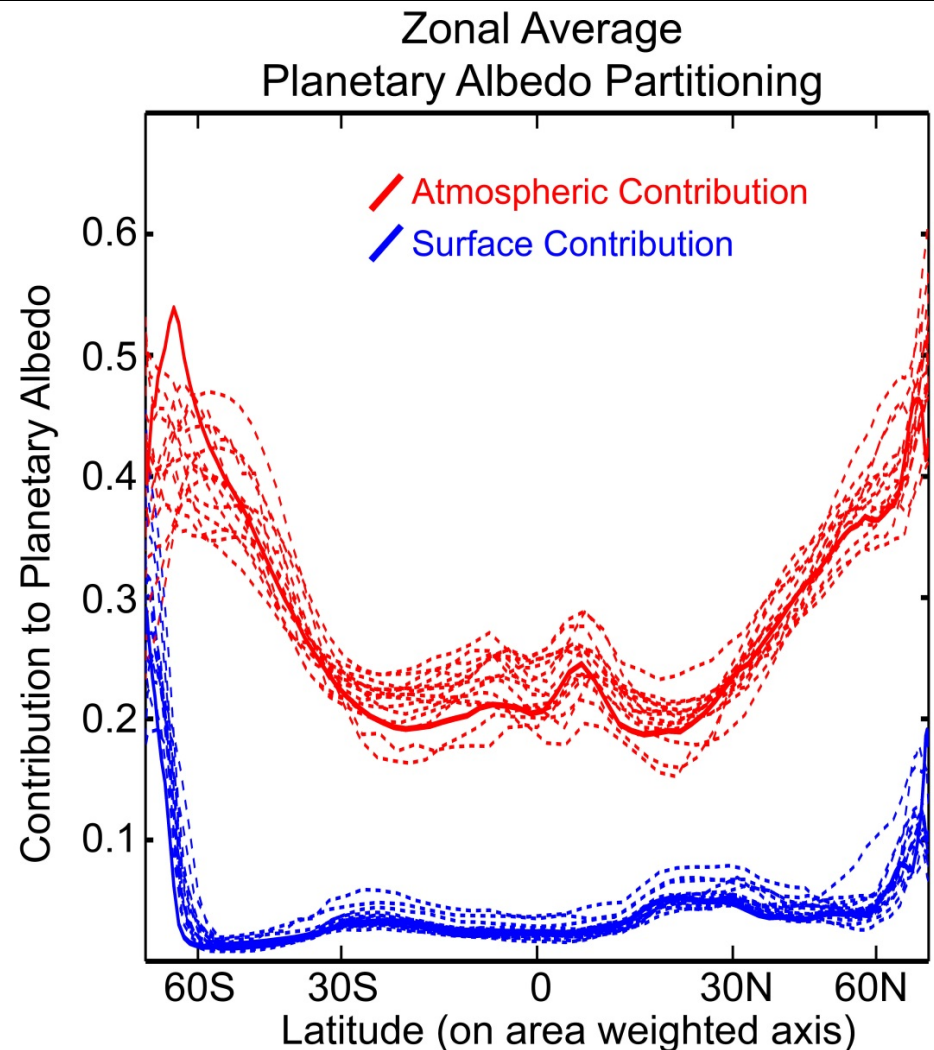
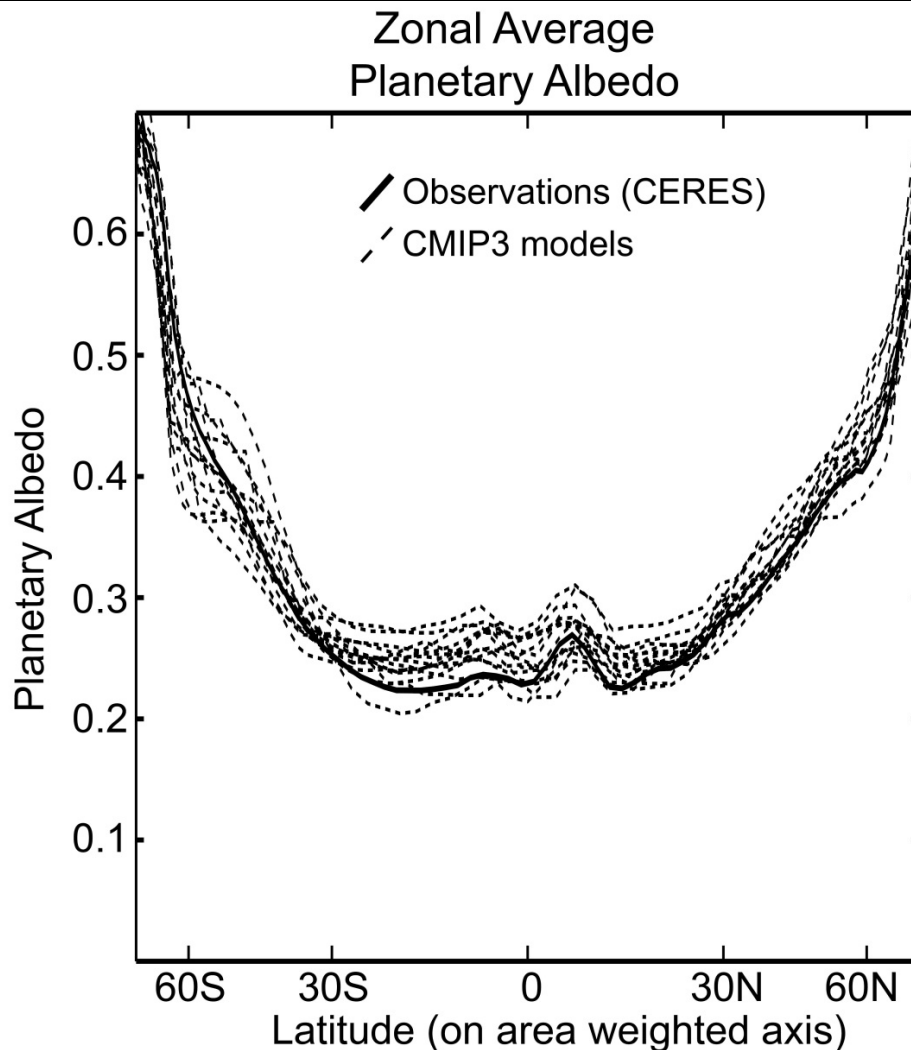
# What determines the equator-to-pole contrast of planetary albedo?



# Surface and atmospheric contributions to the equator-to-pole contrast in absorbed shortwave radiation

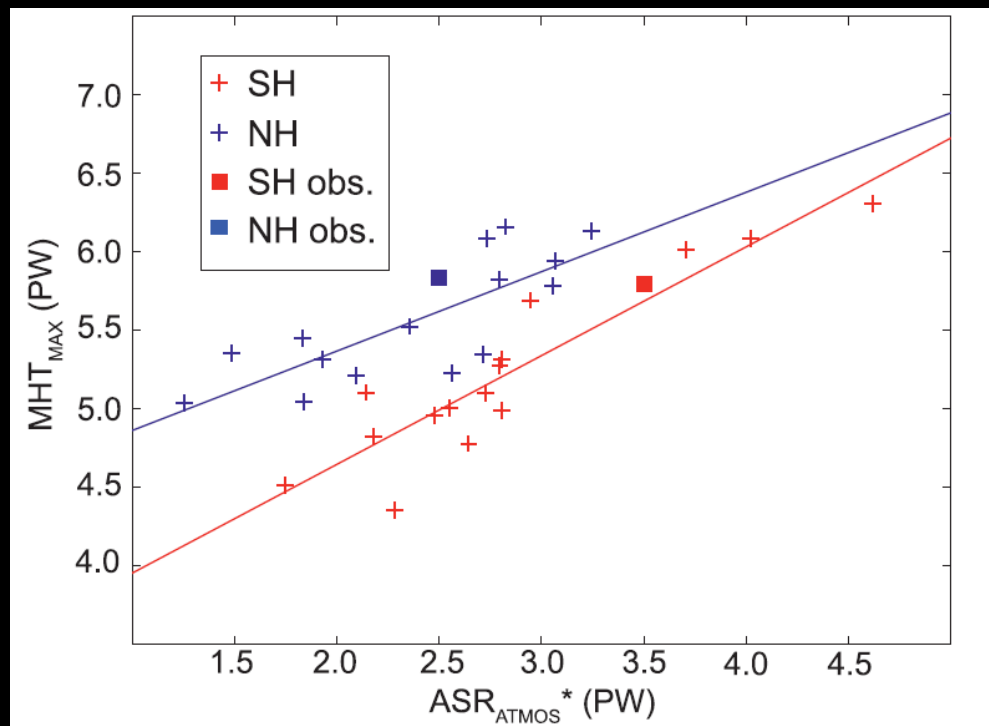


# Planetary Albedo Partitioning



# Atmospheric and Surface reflection contribution to ASR\*

Poleward Energy Transport



Equator to pole contrast in clouds

Cloud Reflection



Equator to pole  
contrast of absorbed  
shortwave



Poleward Energy  
Transport

### 3. The seasonal cycle of atmospheric heating and temperature

30% reflected

20% absorbed  
atmosphere

50% transmitted  
To surface

50% net flux  
From surface  
To  
atmosphere

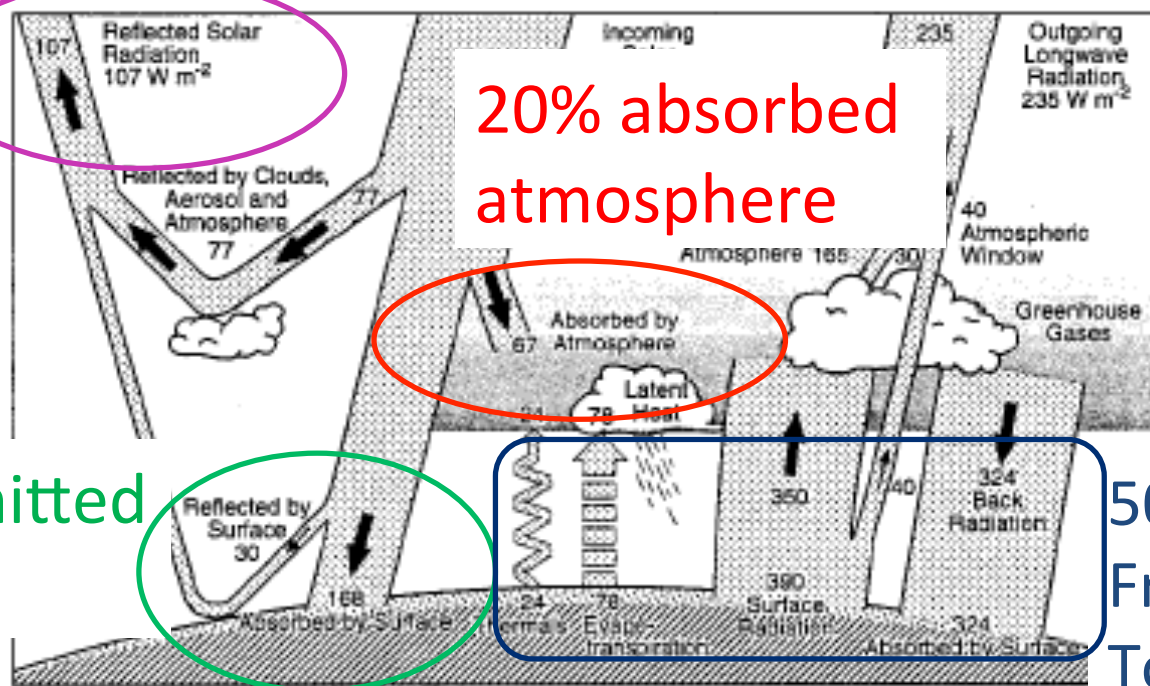
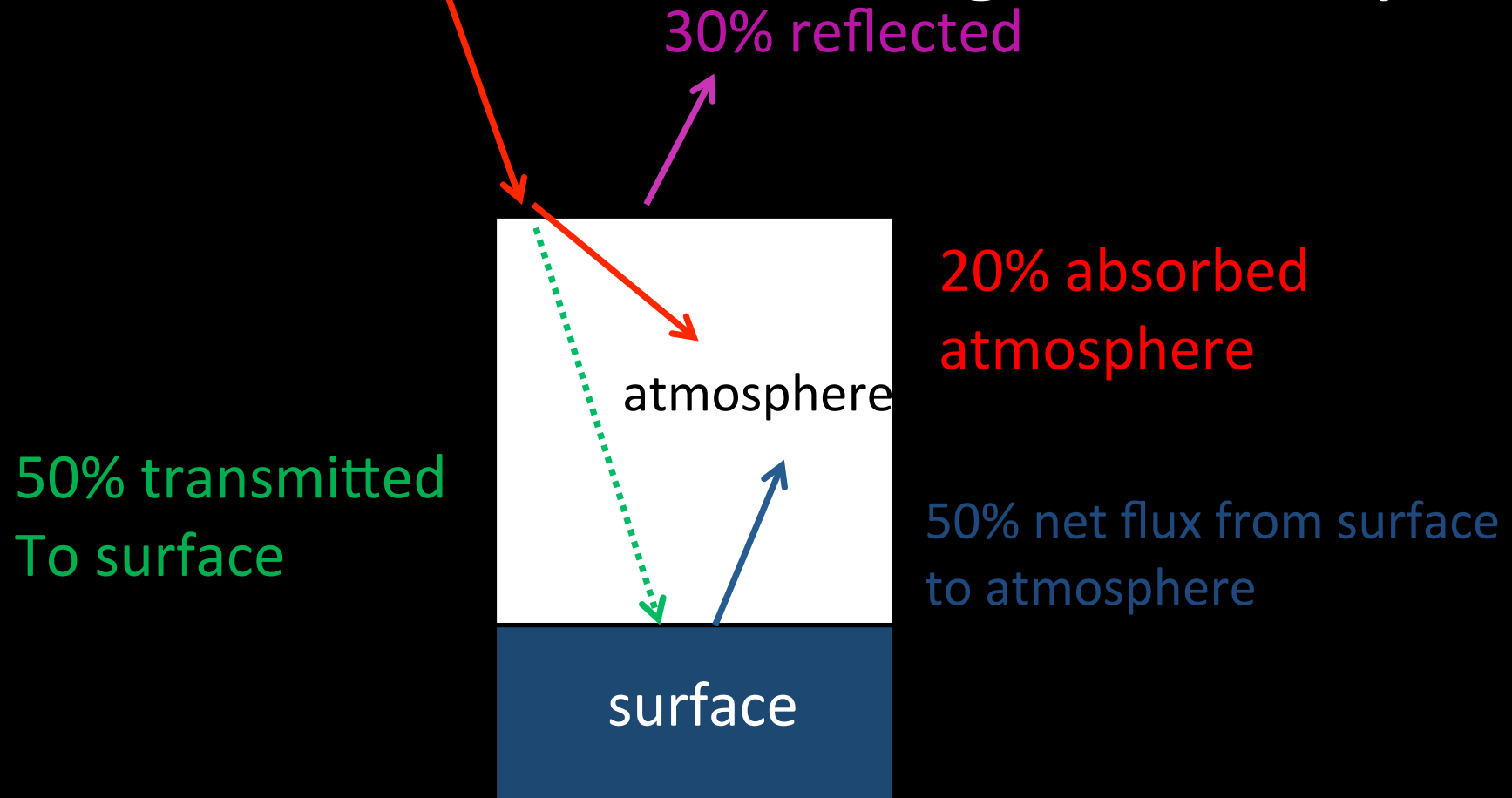


FIG. 7. The earth's annual global mean energy budget based on the present study. Units are  $\text{W m}^{-2}$ .

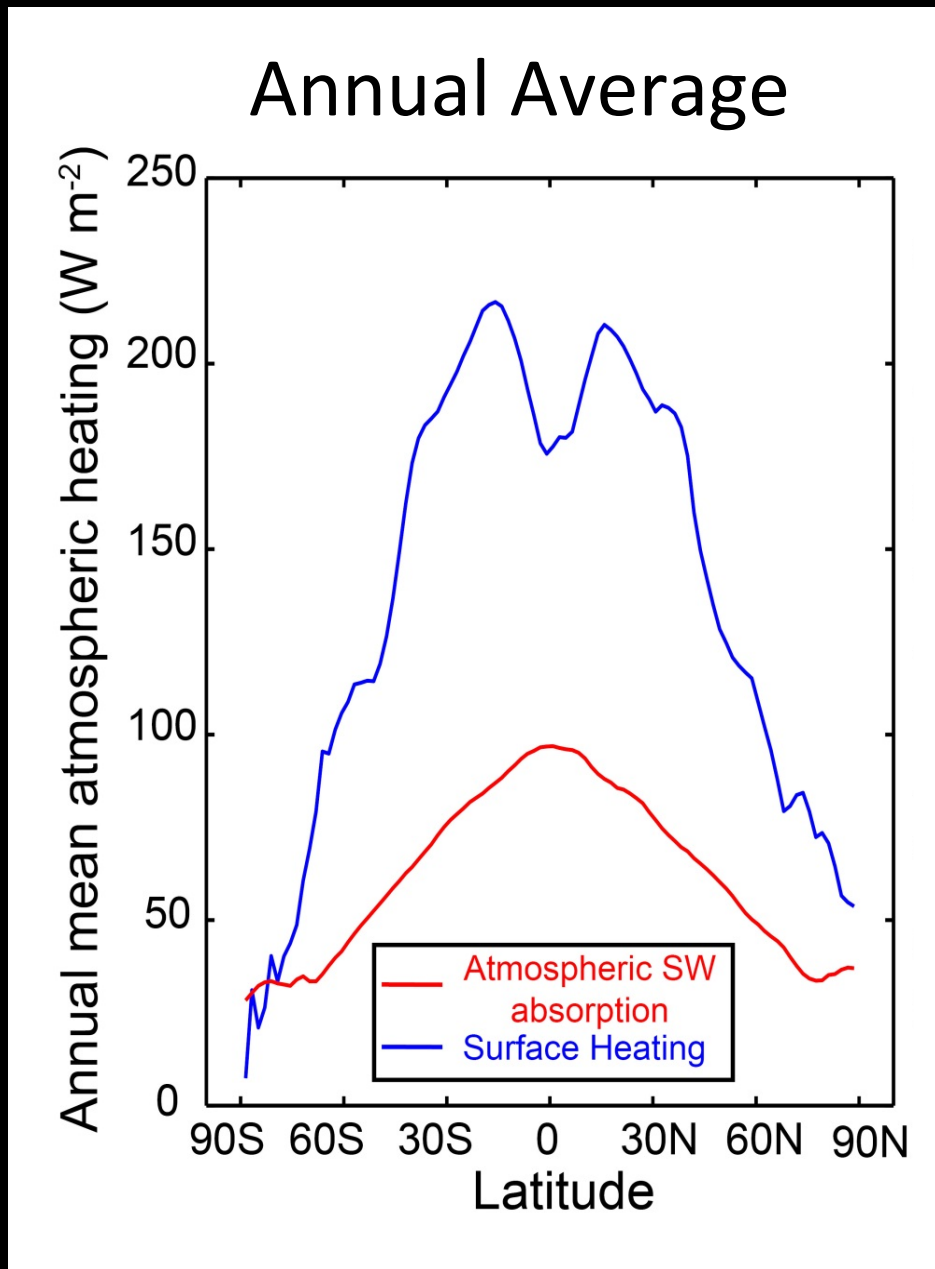
# Annual mean heating summary



Surface energy balance (absorbed solar flux = upward flux to the atmosphere) requires that the ratio of heating from direct absorption to surface energy fluxes is approximately:

atmospheric absorption (SW) /atmospheric transmissivity(SW)  
=> direct absorption accounts for 30% of atmospheric heating

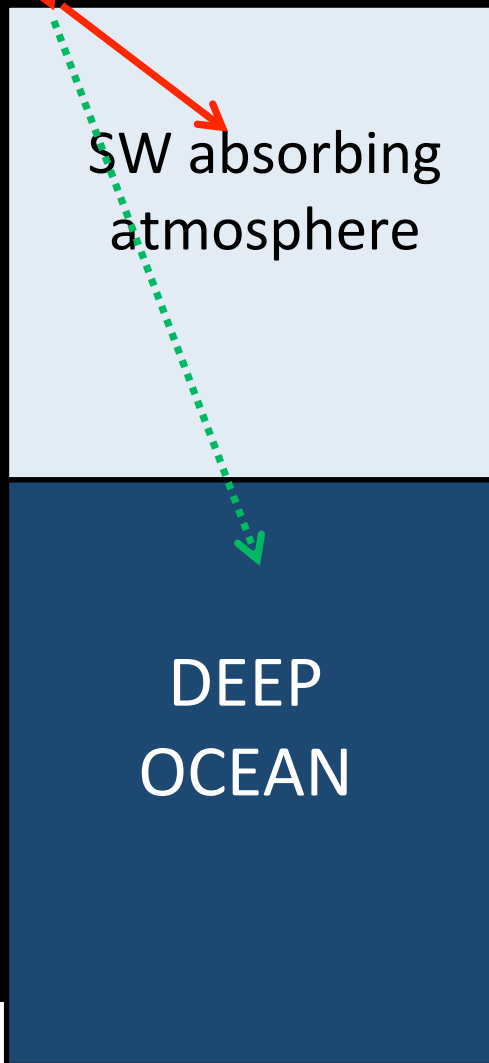
# Spatial structure of atmospheric heating



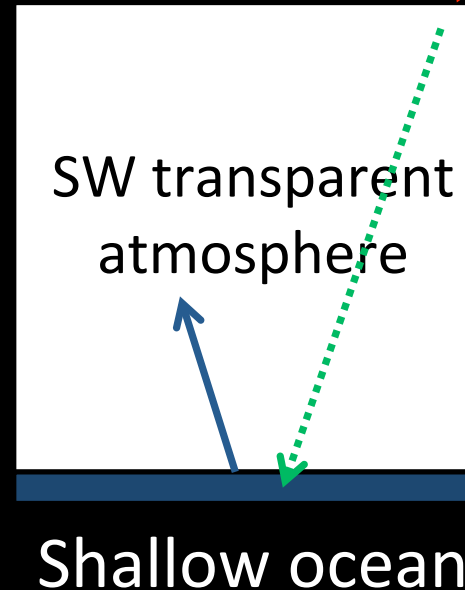
# Limiting models of seasonal atmospheric heating

— SW flux  
— Turbulent + LW  
Surface flux

Heated from above



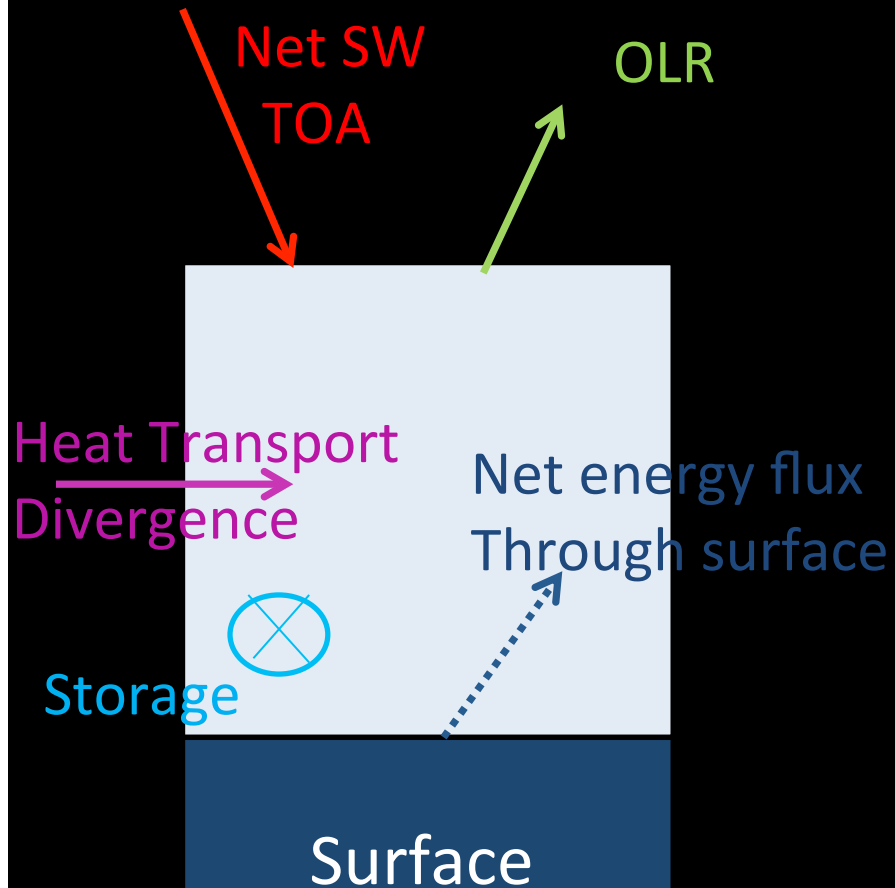
Heated from below





# Atmospheric energy budget

## Conventional



Net energy flux through surface =  
Solar + turbulent + LW

# Atmospheric energy budget -- Observations

$$\frac{1}{g} \int_0^{P_s} \frac{dE}{dt} dP = SWABS + SHF - OLR - \frac{1}{g} \int_0^{P_s} \nabla \cdot (\vec{U} E) dP$$

Storage

Heat Transport Divergence

$$SWABS = SW \downarrow_{TOA} - SW \uparrow_{TOA} + SW \uparrow_{SURF} - SW \downarrow_{SURF}$$

$$SHF = SENS \uparrow_{SURF} + LH \uparrow_{SURF} + LW \uparrow_{SURF} - LW \downarrow_{SURF}$$

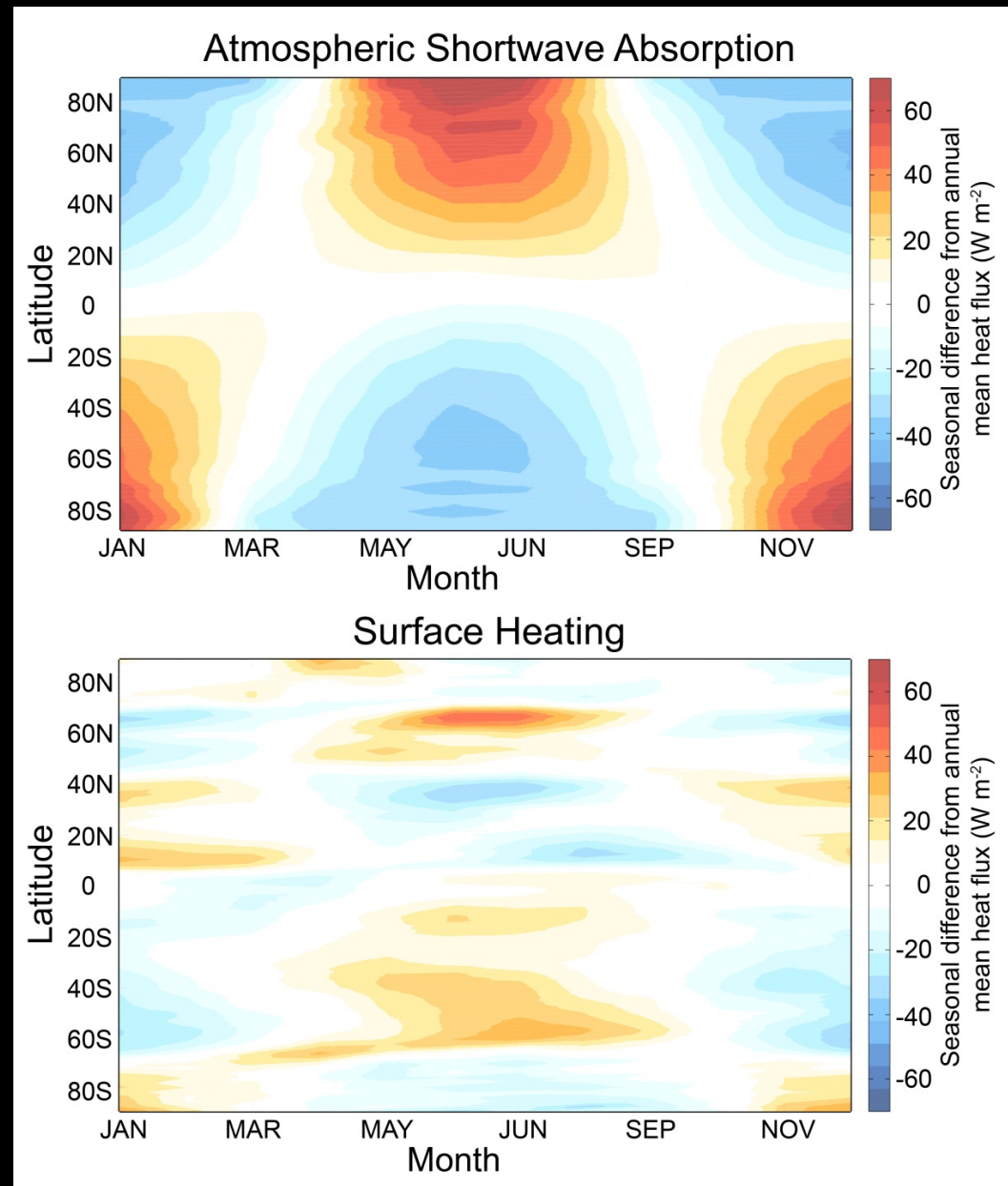
## Radiative Fluxes from CERES

Dynamic fluxes and storage are calculated from ERA-40 reanalysis

Atmos. Surface exchange (SHF) as RESIDUAL

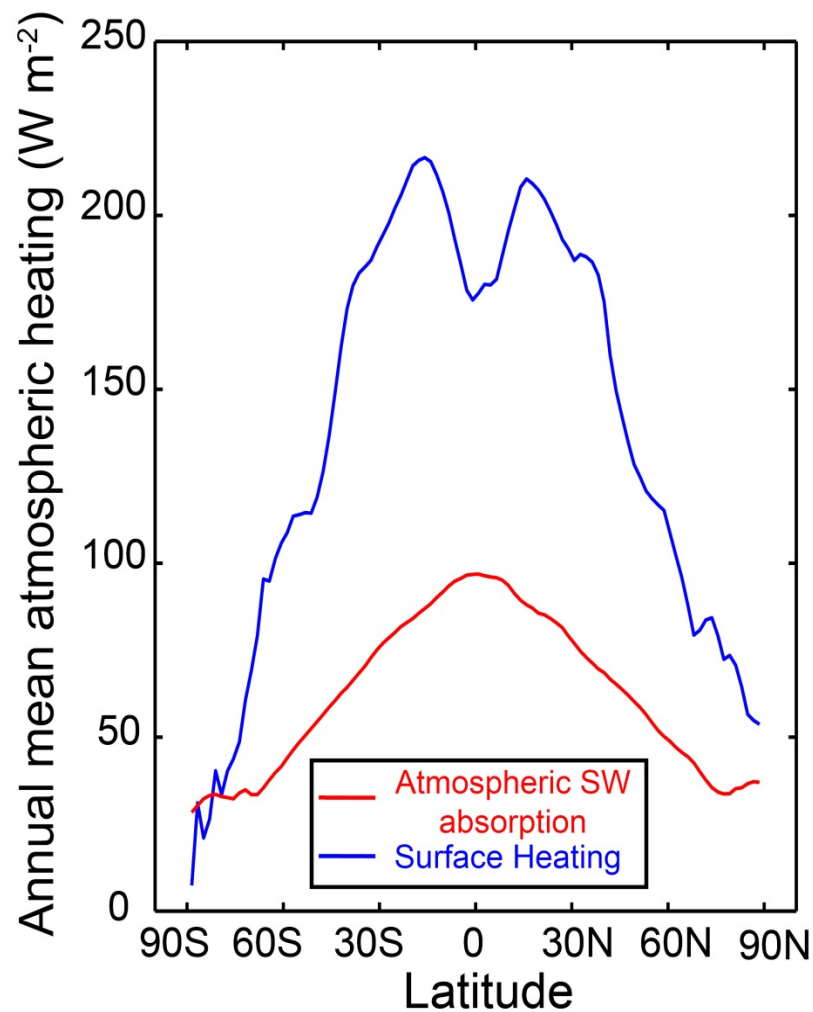
# Seasonal Atmospheric Heating

ANNUAL  
MEAN IS  
REMOVED

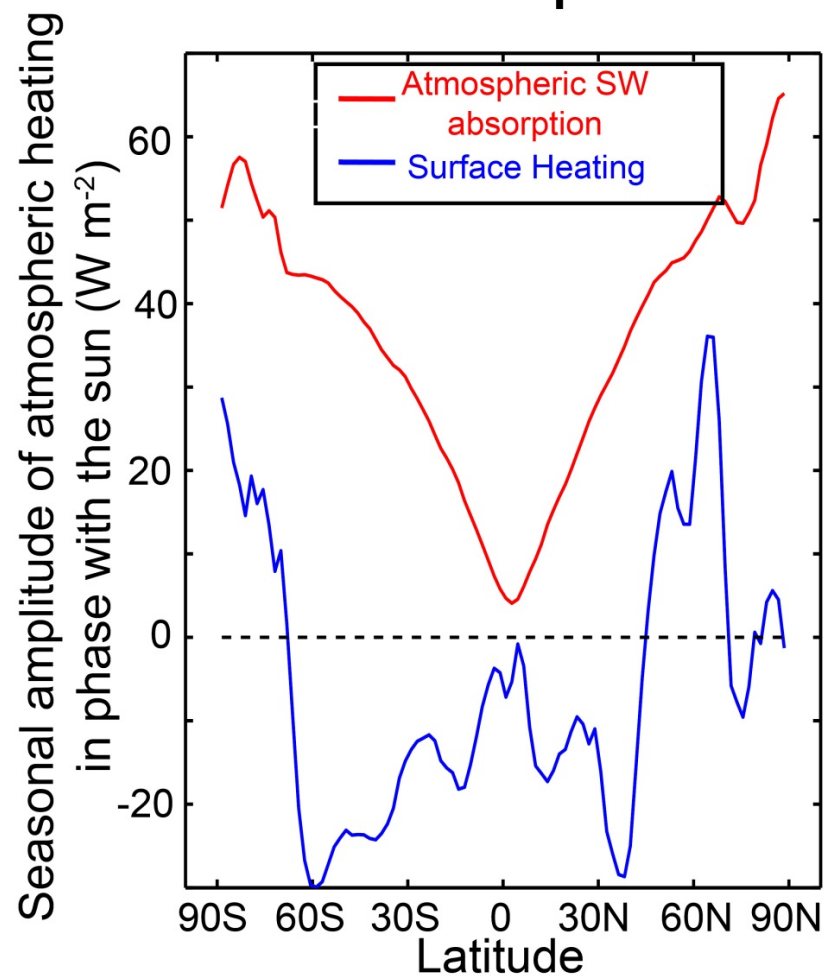


# Seasonal Amplitude of heating

## Annual Mean



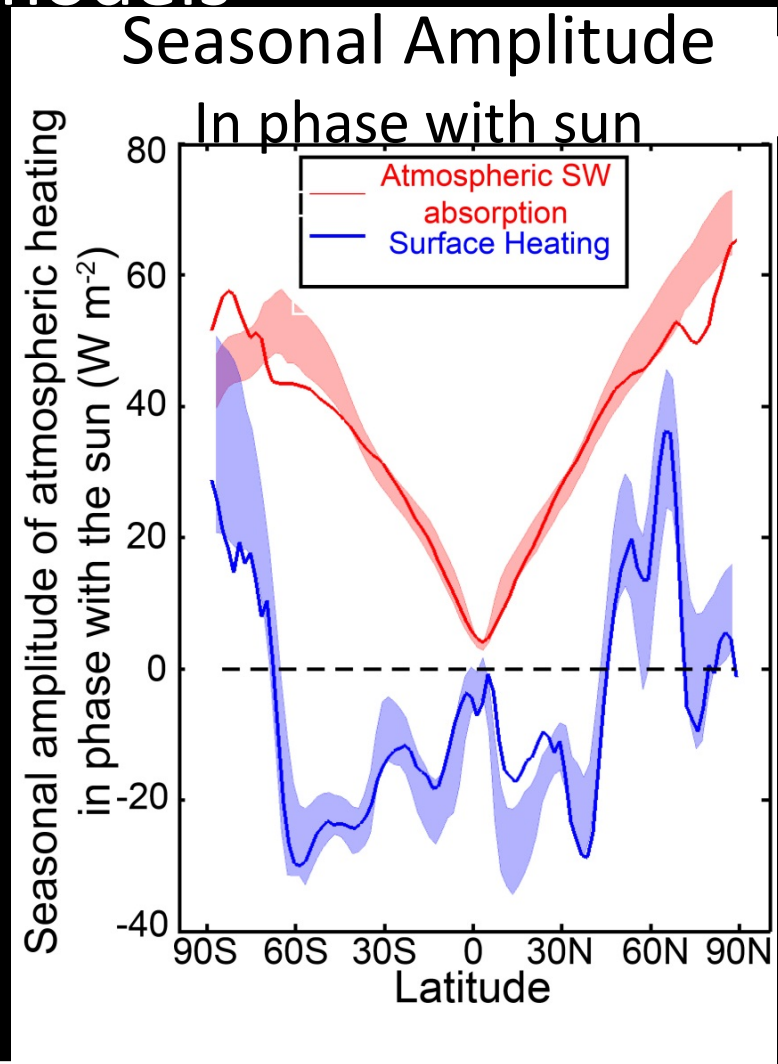
## Seasonal Amplitude



# The seasonal heating of the atmosphere in climate models

Surface fluxes and shortwave absorption are calculated directly from model output

Atmospheric heat transport as a residual

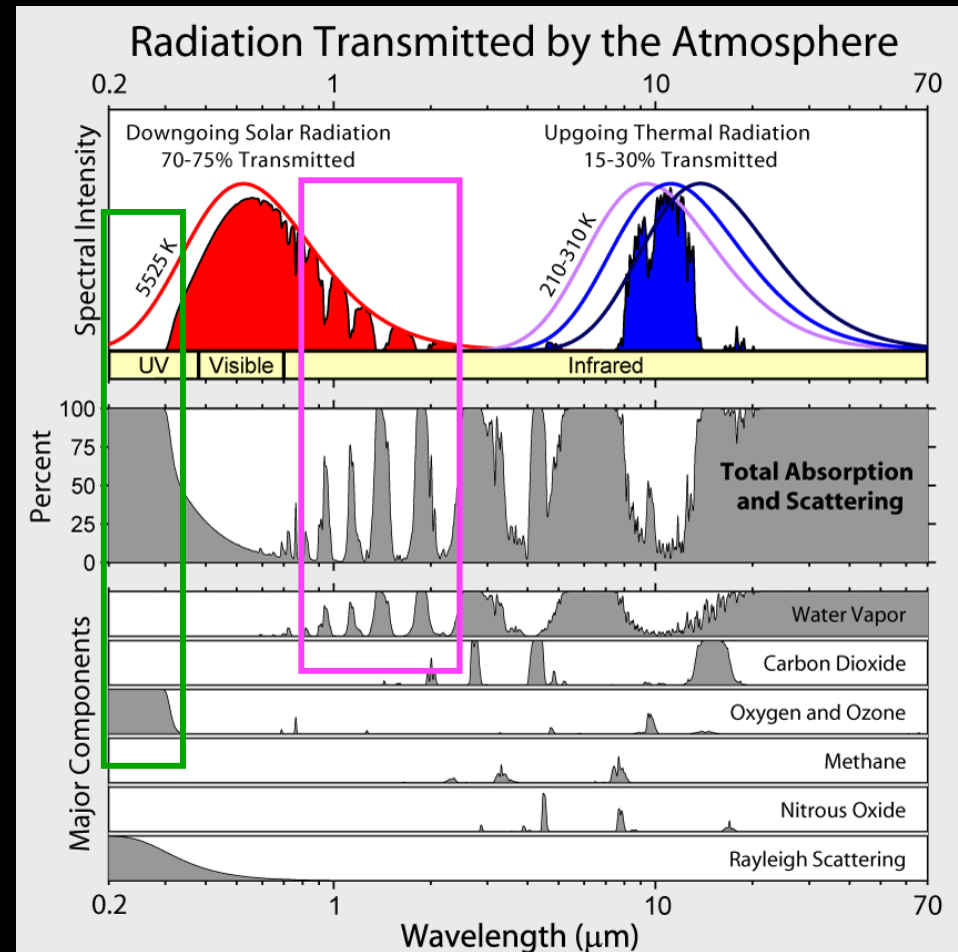
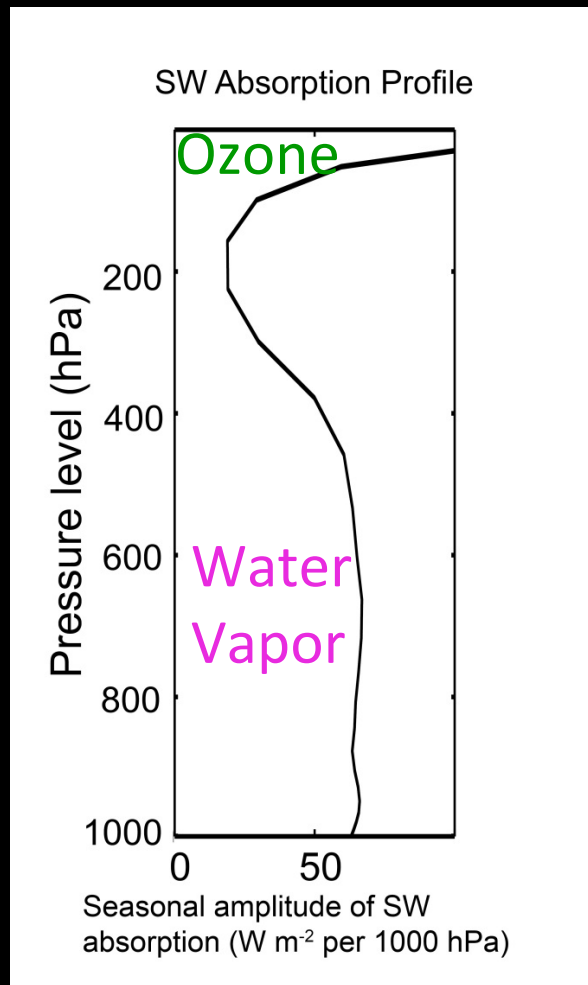


=> Solid lines = observations

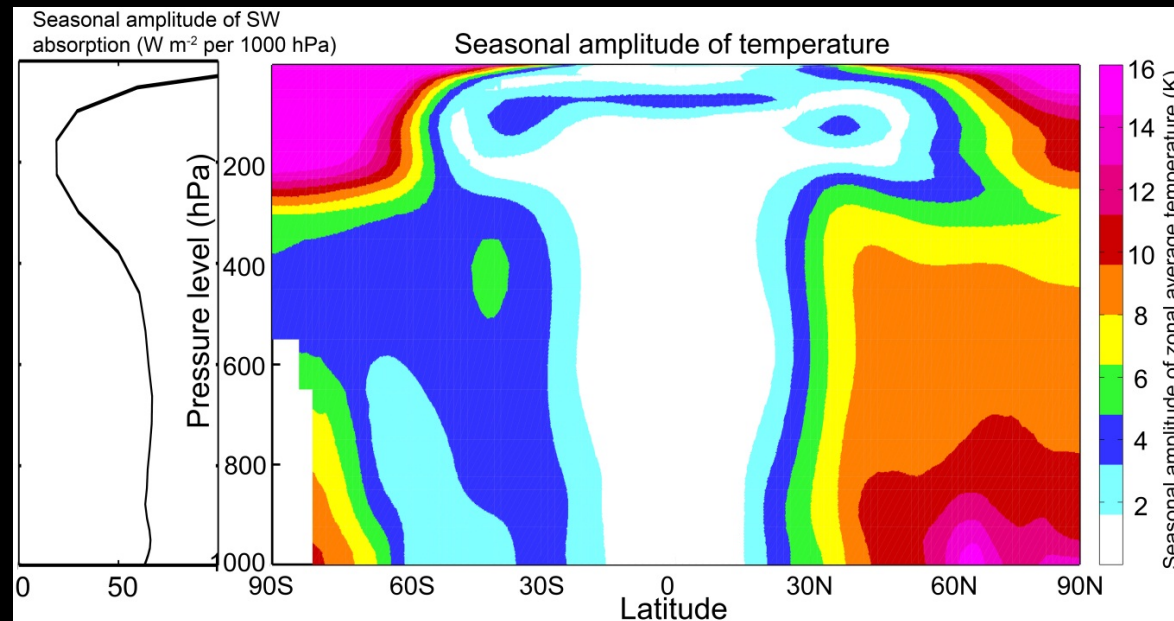
=> Shading is  $\pm 1\sigma$  about CMIP3 ensemble mean

# Vertical structure of SW absorption

GFDL Model  
Seasonal amplitude  
In the extratropics



# Structure of seasonal amplitude in temperature

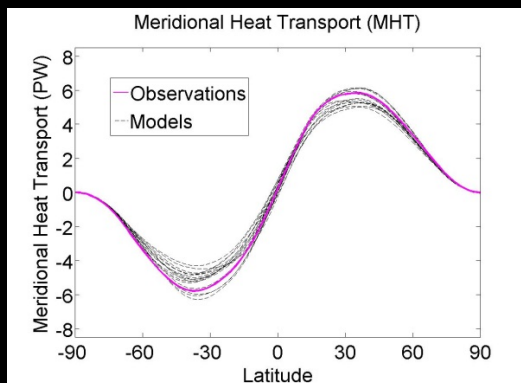
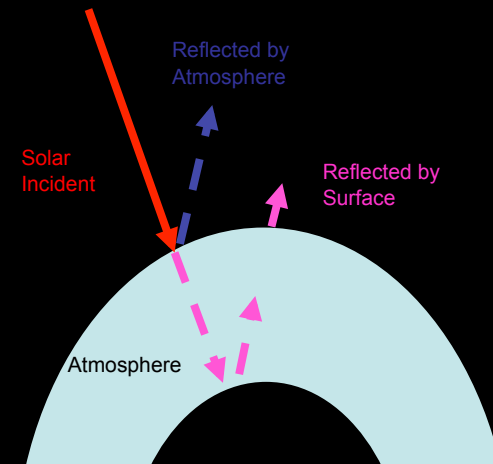


Seasonal cycle is surface amplified where the surface heat fluxes contribute to seasonal heating (over land)



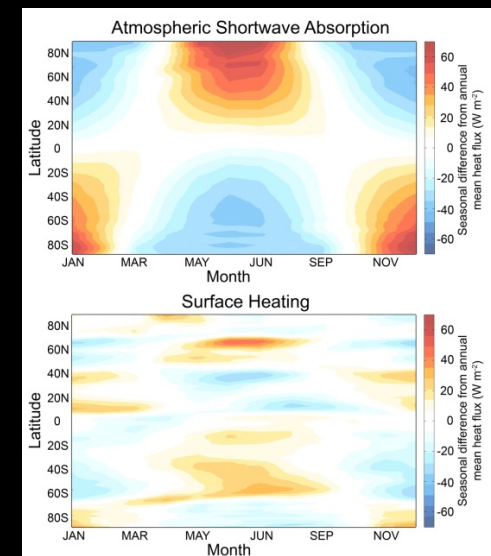
# Conclusions

1. Global mean planetary albedo is primarily (88%) due to atmospheric reflection and only secondarily (12%) due to surface reflection
  - > Climatology and model spread
  - > Hemispheric contrast in planetary albedo sets ITCZ location



2. Poleward energy transport varies by 20% in climate models and is a consequence of simulated clouds

3. The seasonal heating of the atmosphere is due to shortwave heating of the atmosphere and is opposed by surface fluxes (contrast to annual mean)

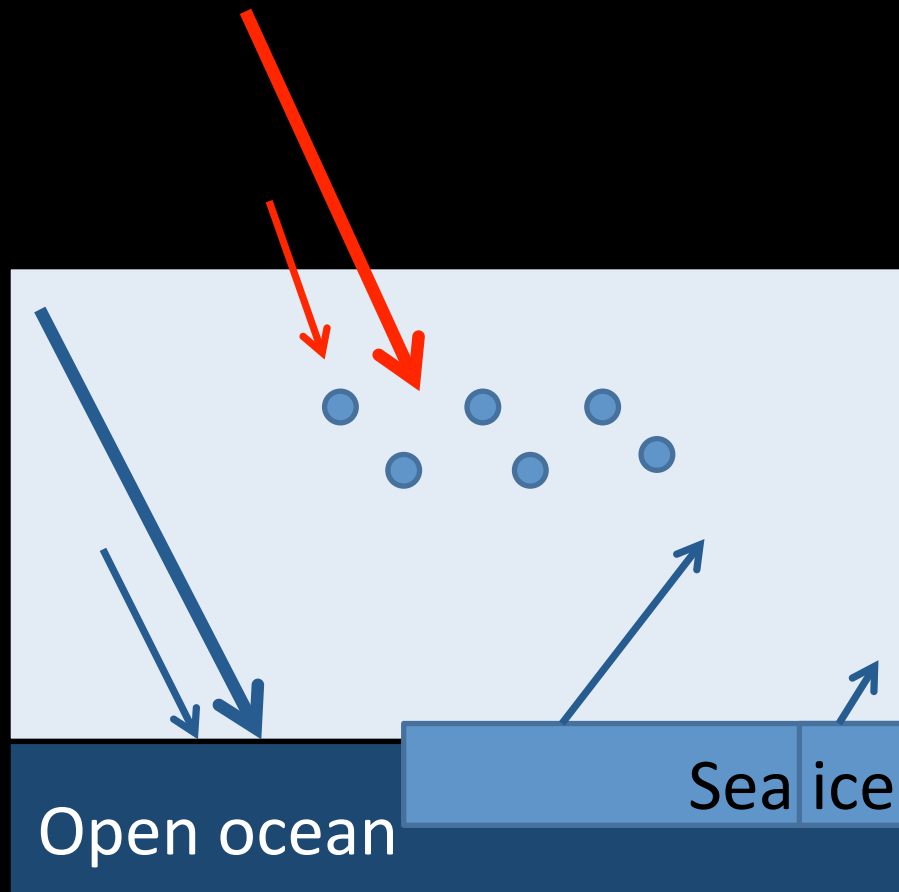




# Extras

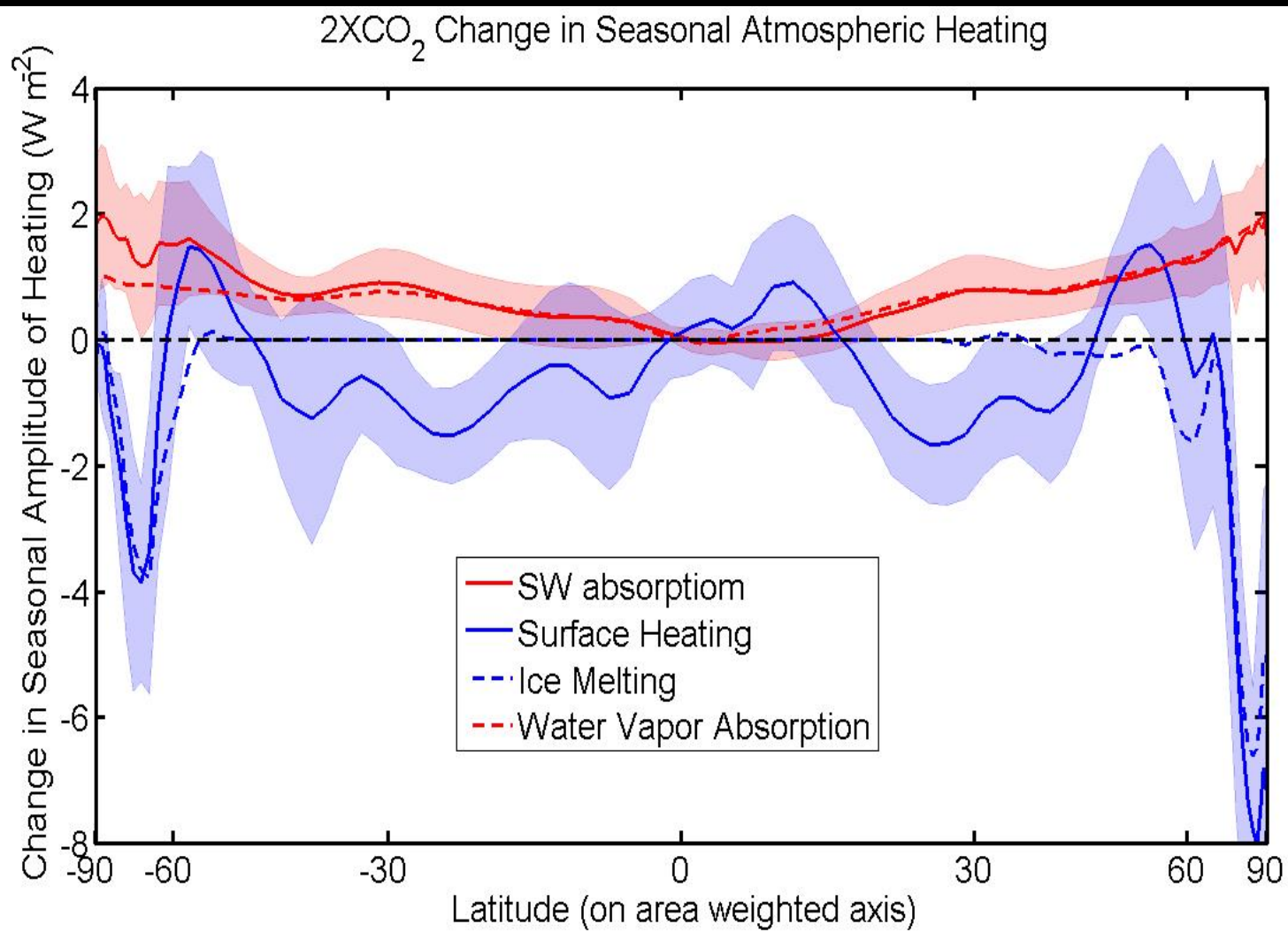
# CO<sub>2</sub> doubling expectations

## Summer schematic

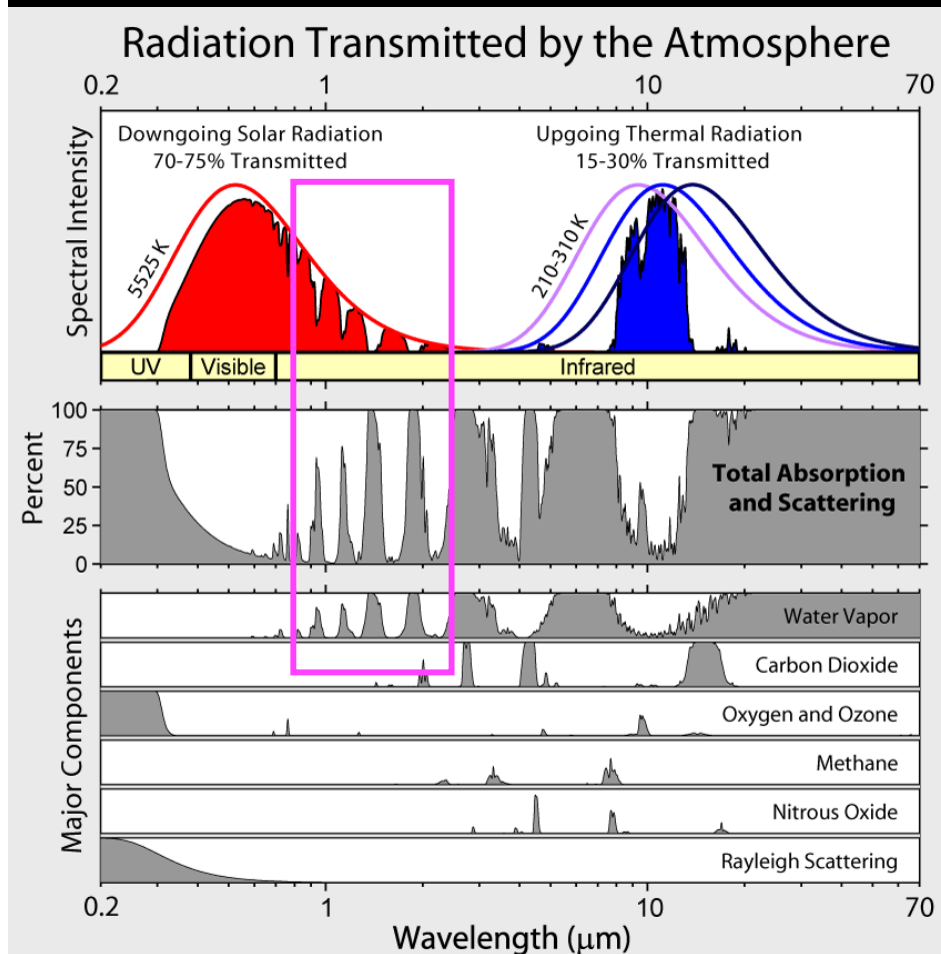


Surface flux  
SWABS

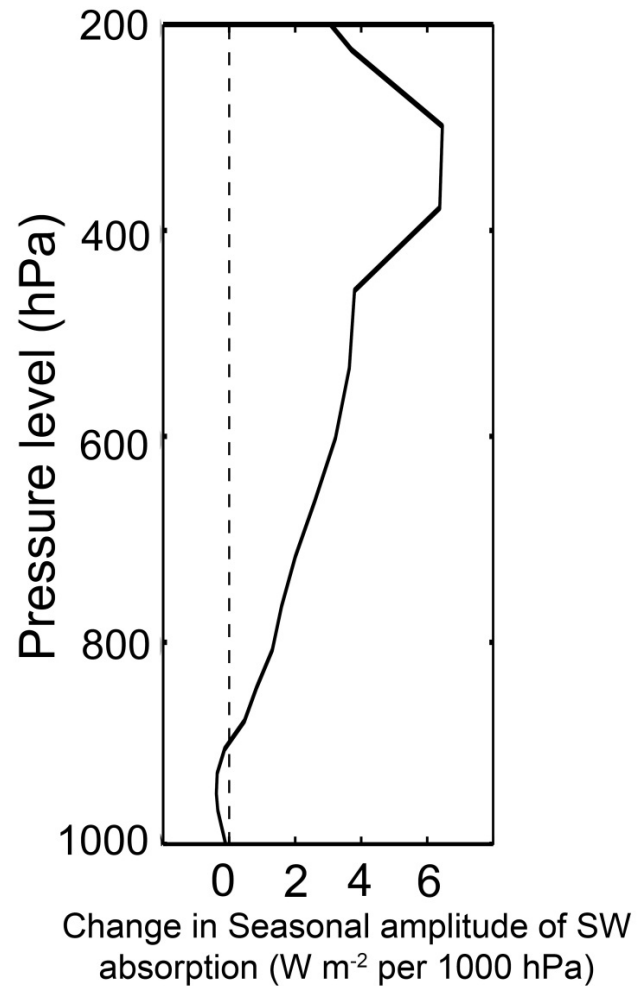
Moistening and melting  
⇒ More seasonal energy input directly into the atmosphere (SWABS)  
⇒ Less seasonal energy input at the surface (SHF)



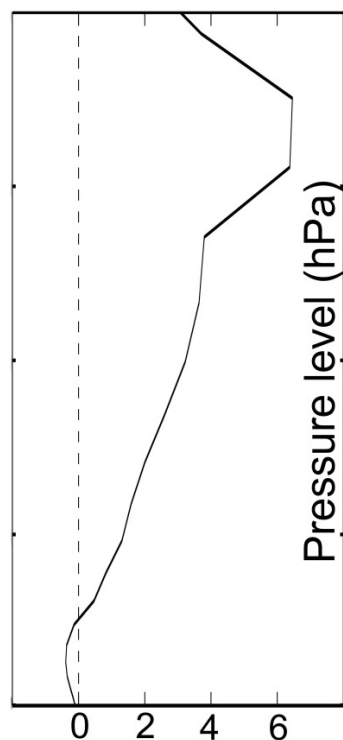
# Water Vapor as a SW Absorber



2XCO<sub>2</sub> Change in SW Absorption Profile in GFDL 2.1

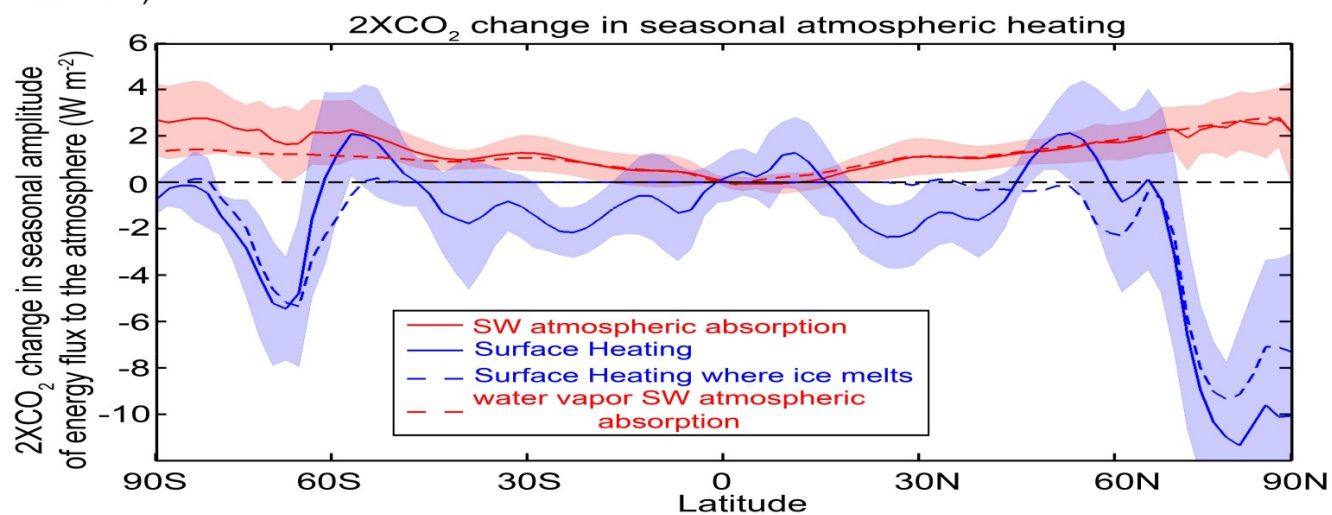
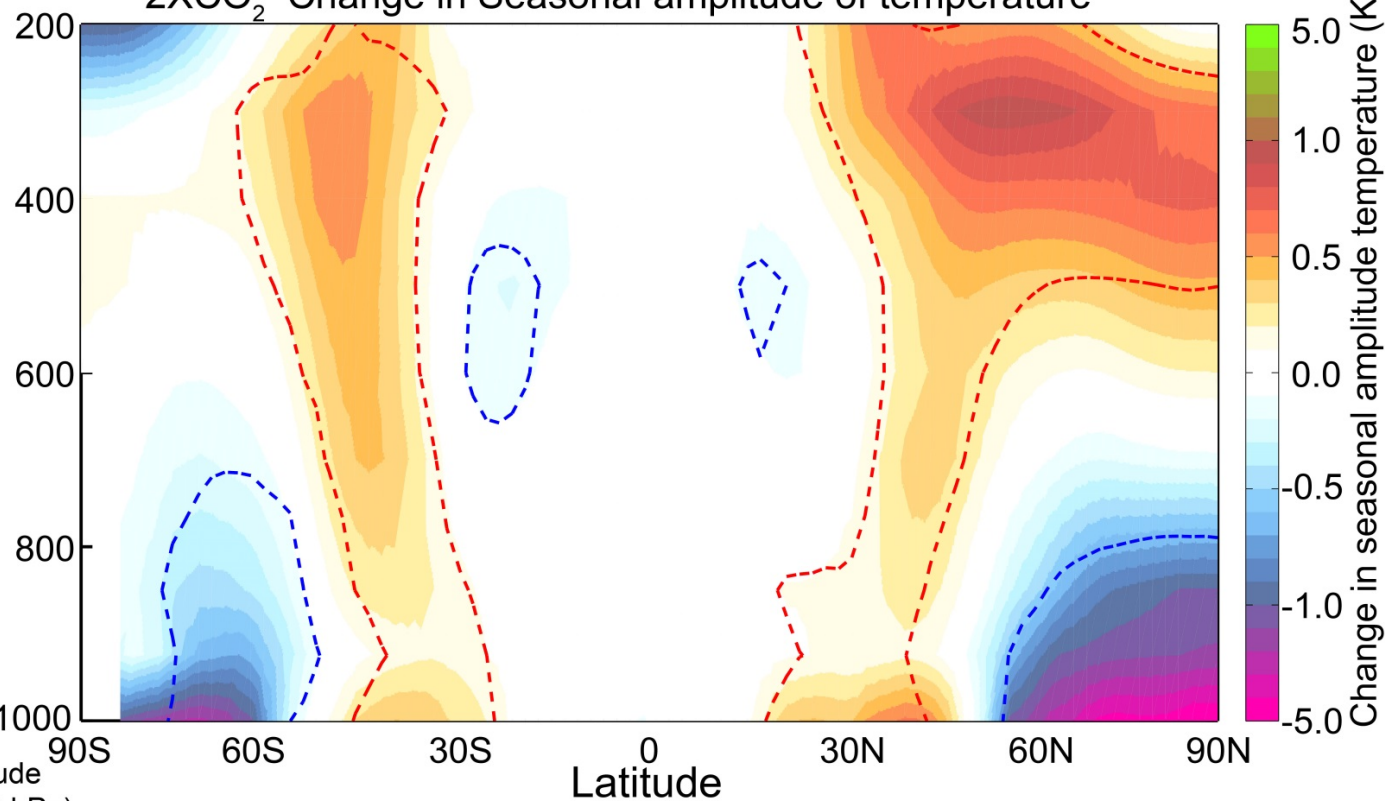


2XCO<sub>2</sub> change SW Abs.

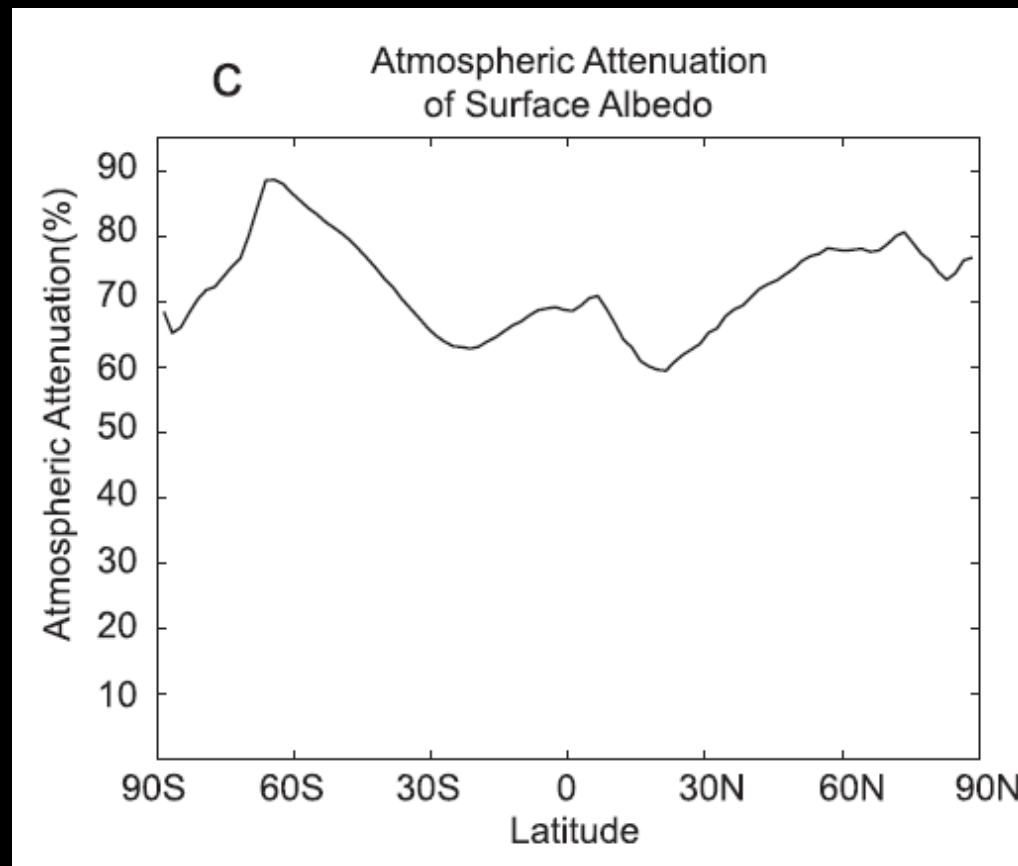


Change in Seasonal amplitude  
of SW abs. (W m<sup>-2</sup> per 1000 hPa)

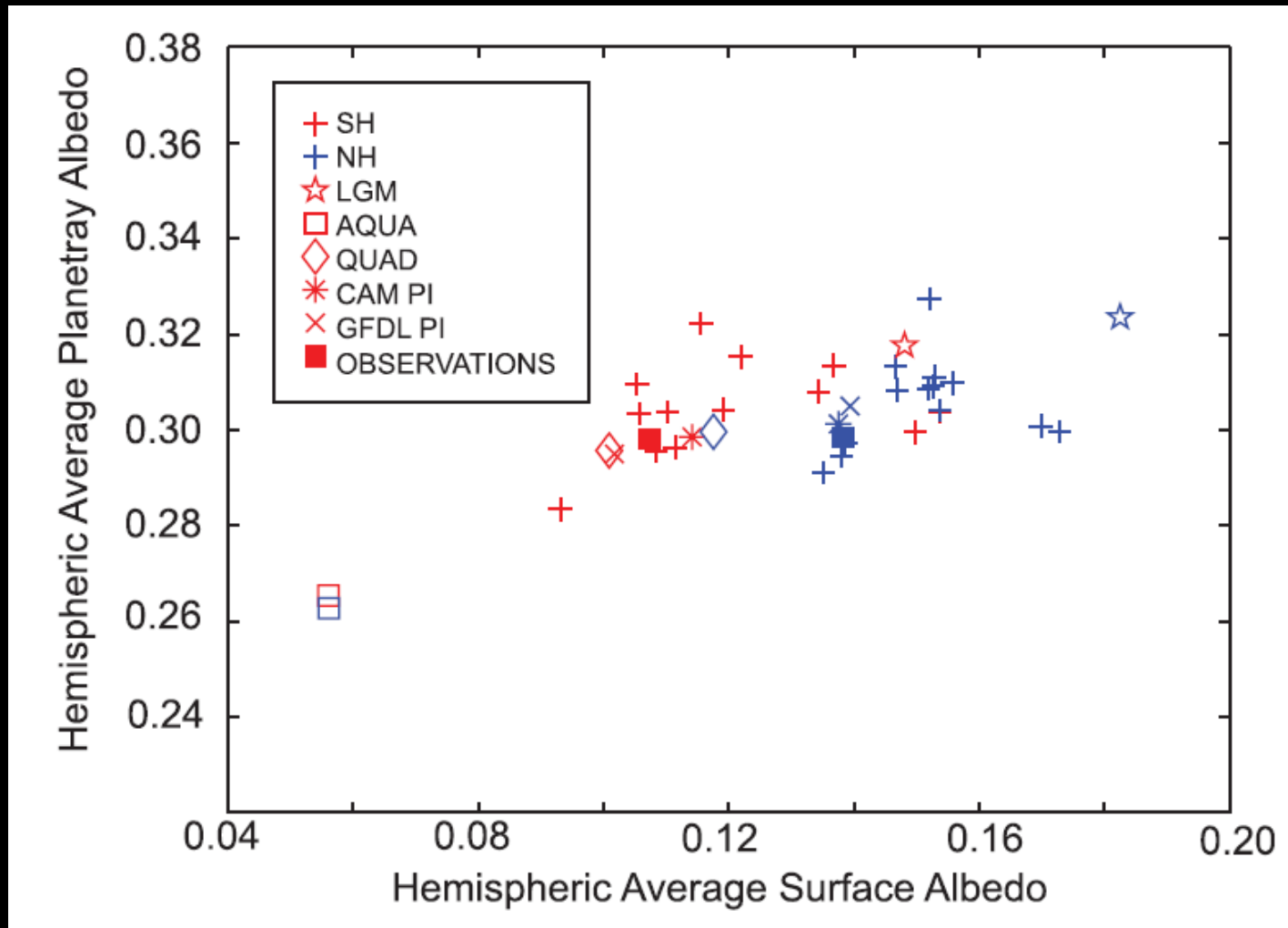
2XCO<sub>2</sub> Change in Seasonal amplitude of temperature



# Meridional structure of atmospheric attenuation



# Planetary Albedo and Surface Albedo



# Meridional profile of albedo in all simulations

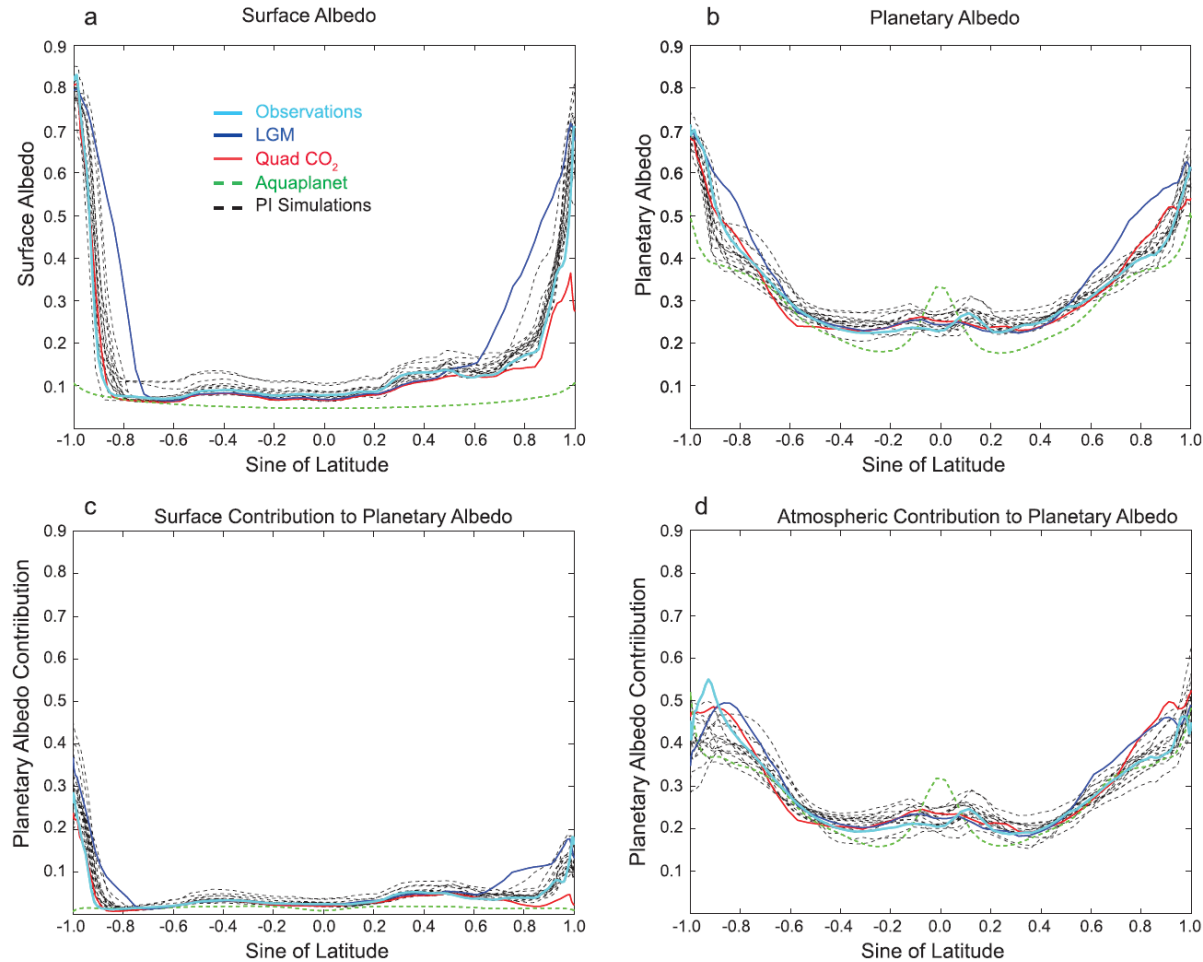
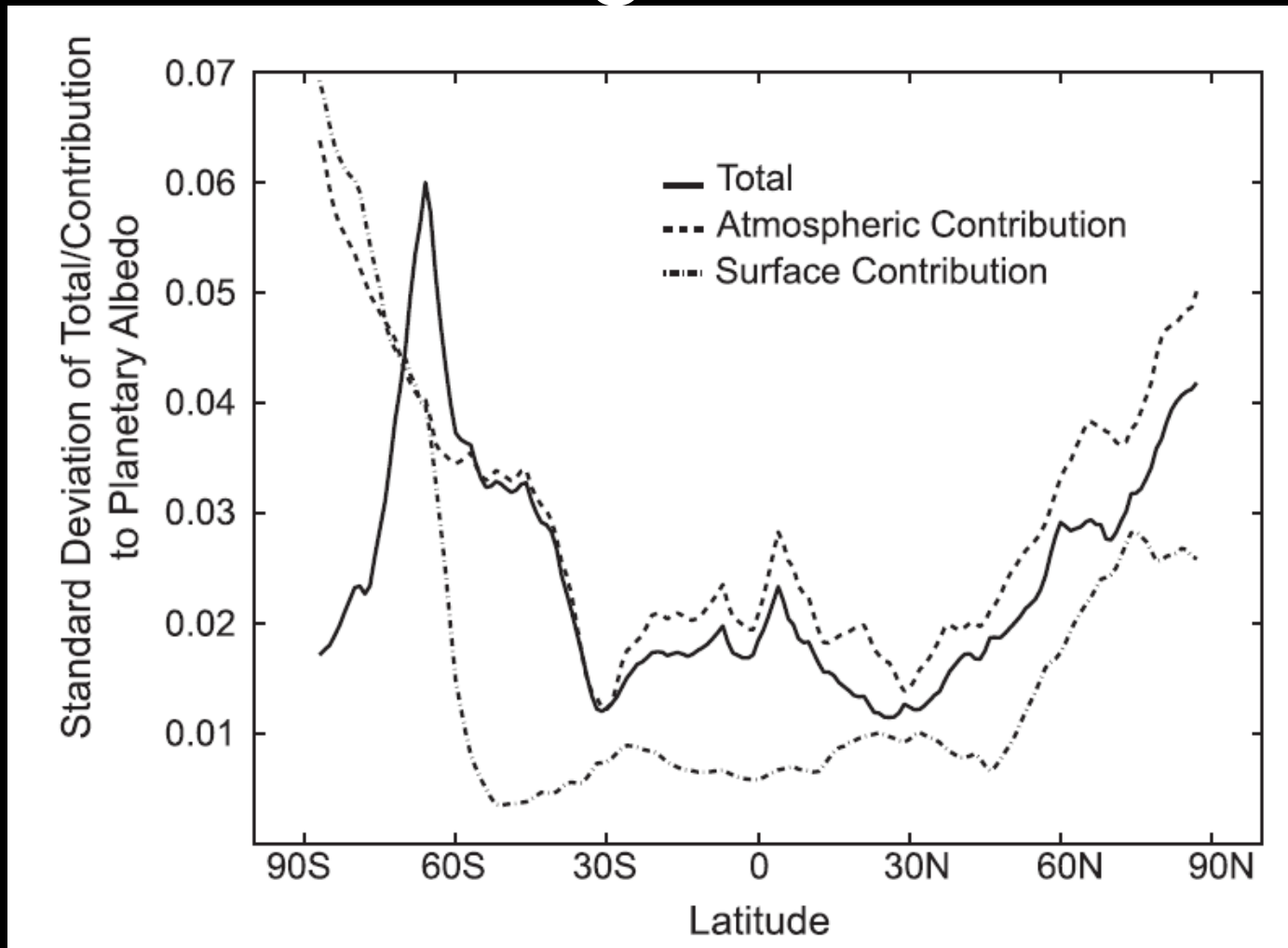


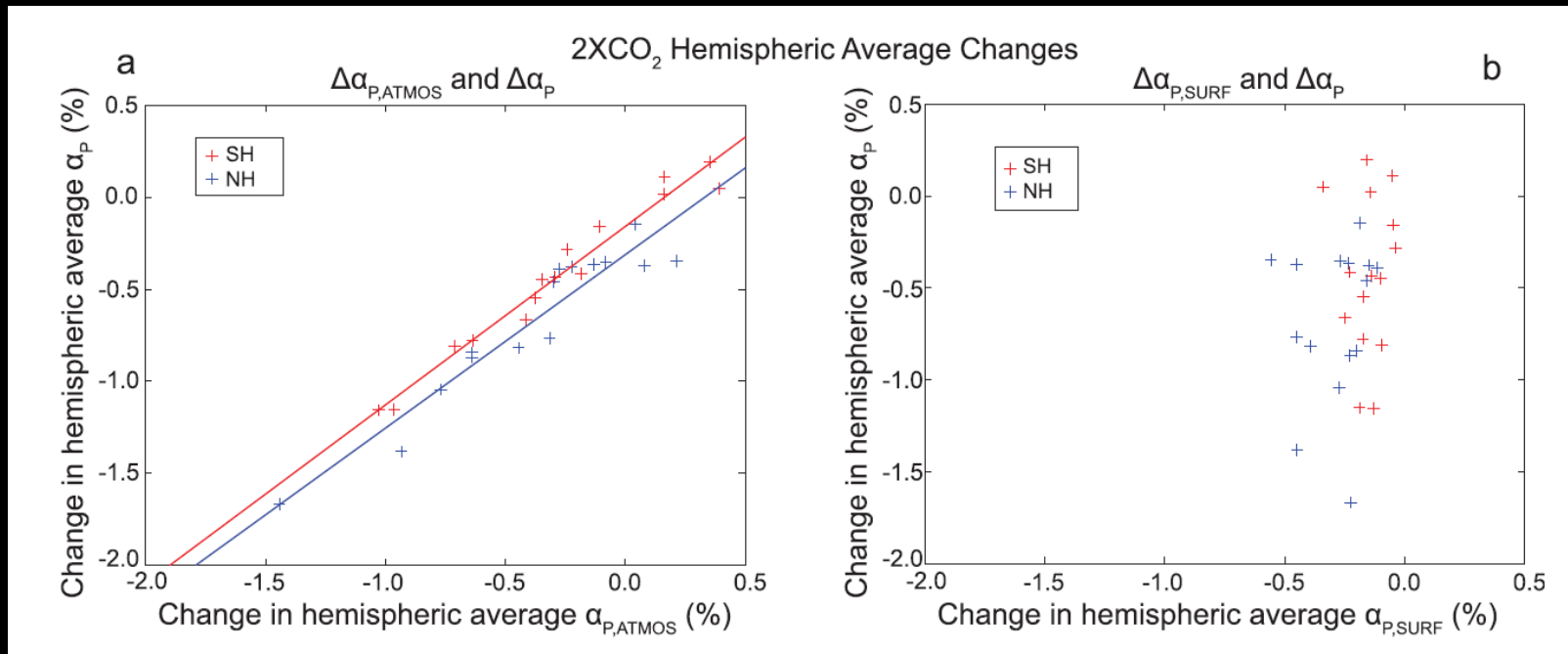
FIG. 6. Zonal annual mean (a)  $\alpha$ , (b)  $\alpha_P$ , (c)  $\alpha_{PSURF}$ , and (d)  $\alpha_{PATMOS}$  in the PI simulations from the CMIP3 models (dashed black lines). Also shown are the observations (solid light blue line) and model simulations of altered climate states (other colored lines).



# Inter-model spread in albedo by region



# 2XCO<sub>2</sub> Planetary Albedo



# 2XCO<sub>2</sub> Planetary Albedo

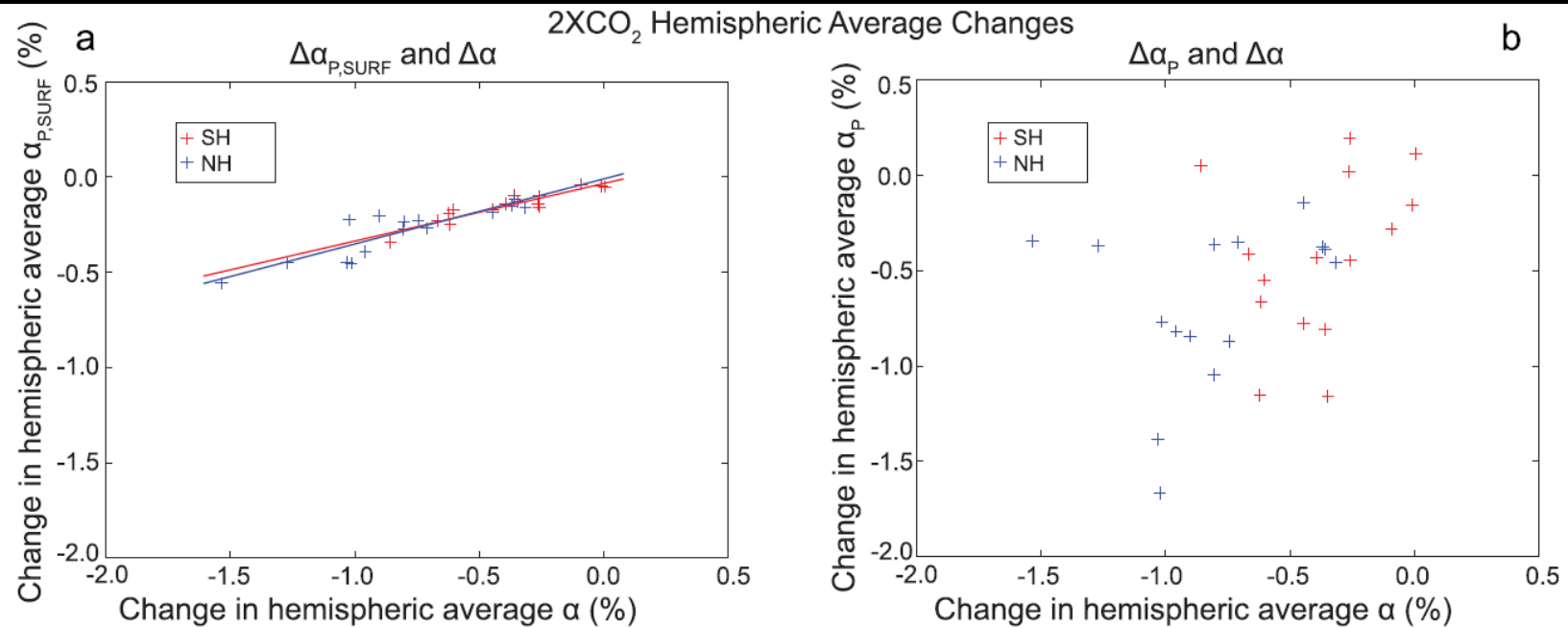
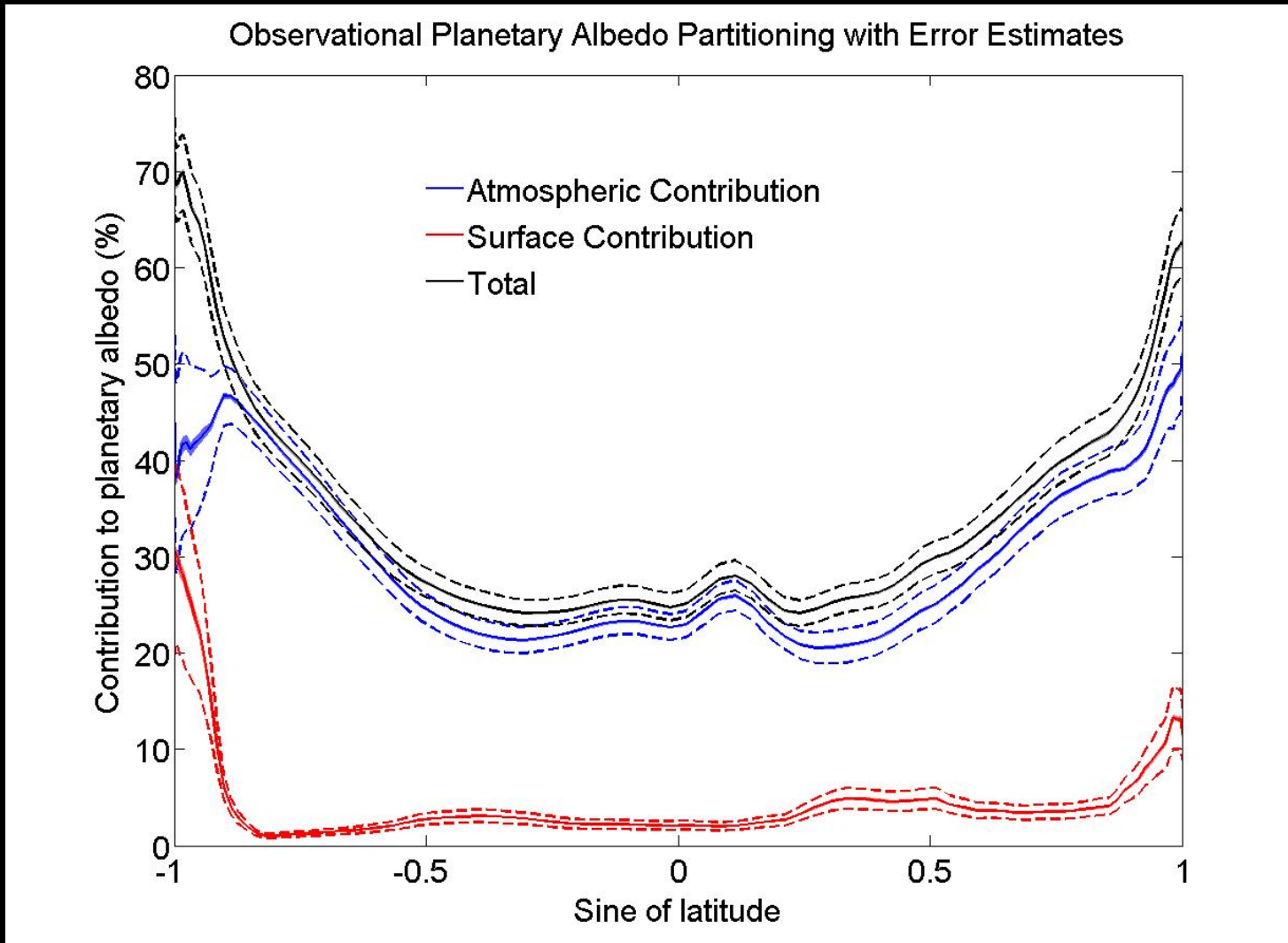
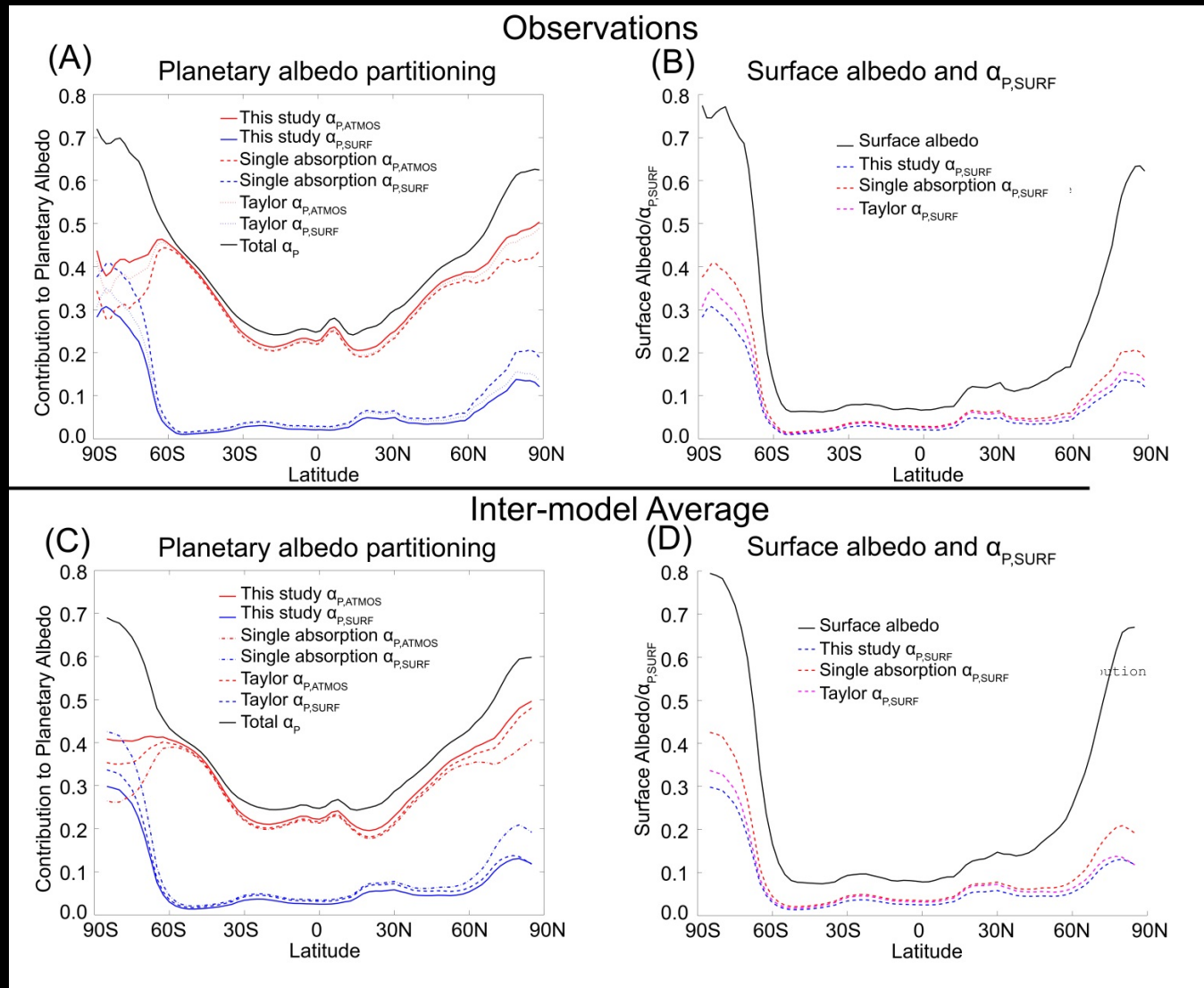


FIG. 9. (a) Change in hemispheric average  $\alpha_{P,SURF}$  in the  $2 \times \text{CO}_2$  runs (relative to the PI simulations) vs change in hemispheric average  $\alpha$ . The lines are the linear best fits in each hemisphere. (b) Change in hemispheric average  $\alpha_P$  in the  $2 \times \text{CO}_2$  runs (relative to the PI simulations) vs change in hemispheric average  $\alpha$ .

# Carl's Question: Planetary Albedo Partitioning Error Bars



# Planetary Albedo Partition: Sensitivity to shortwave absorption assumptions

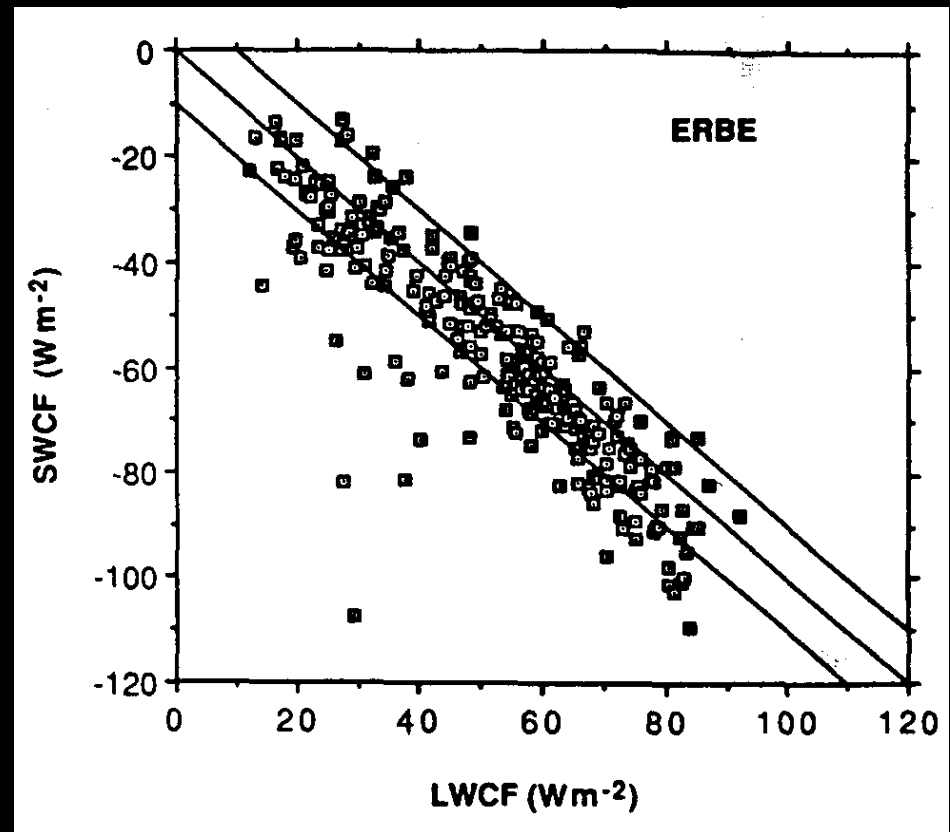


# What determines $OLR^*$ ?

Why don't differences in  $ASR^*$  and  $OLR^*$  compensate for each other?

Longwave Cloud Forcing

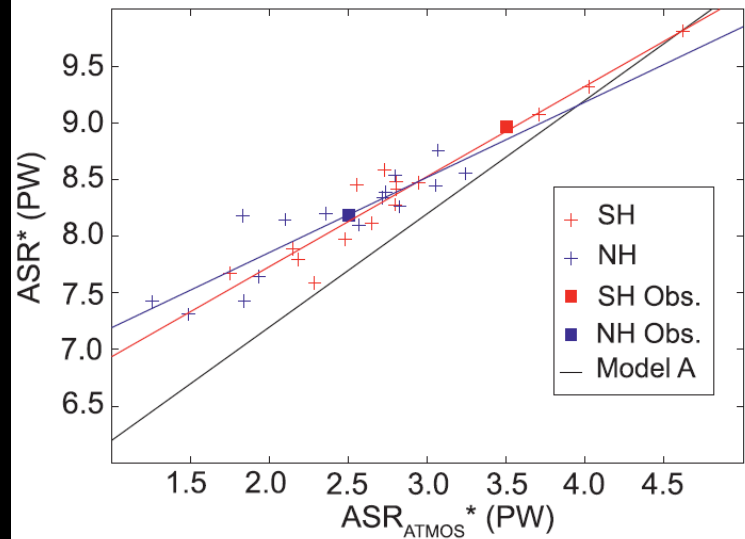
$$LWCF = OLR_{CLEAR} - OLR$$



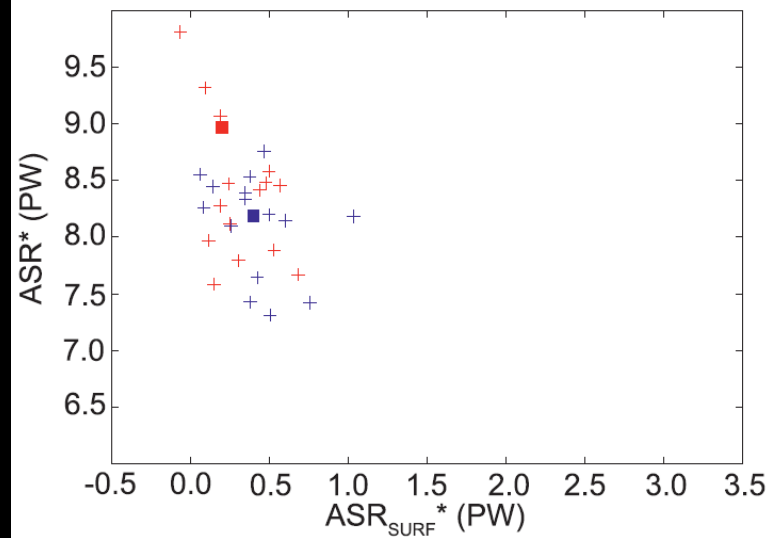
(Kiehl, 1994)

# Contributions to ASR\* spread

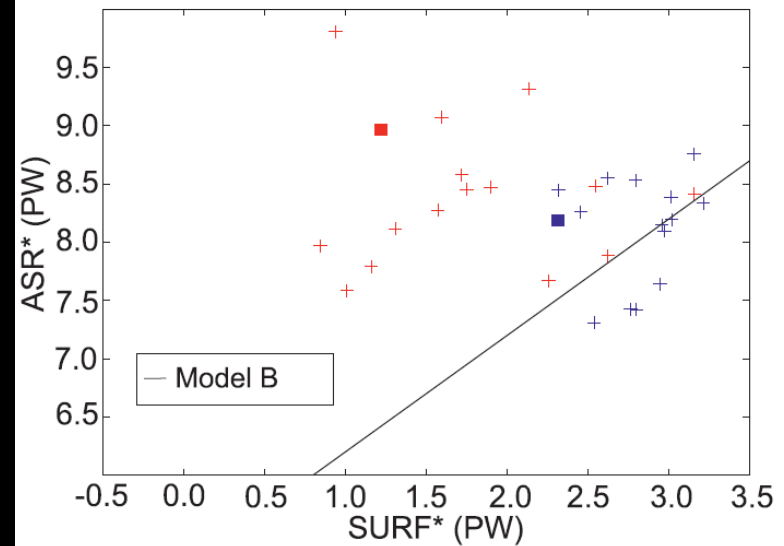
**A** Model Spread in ASR\* and Atmospheric Albedo Contribution to ASR\*



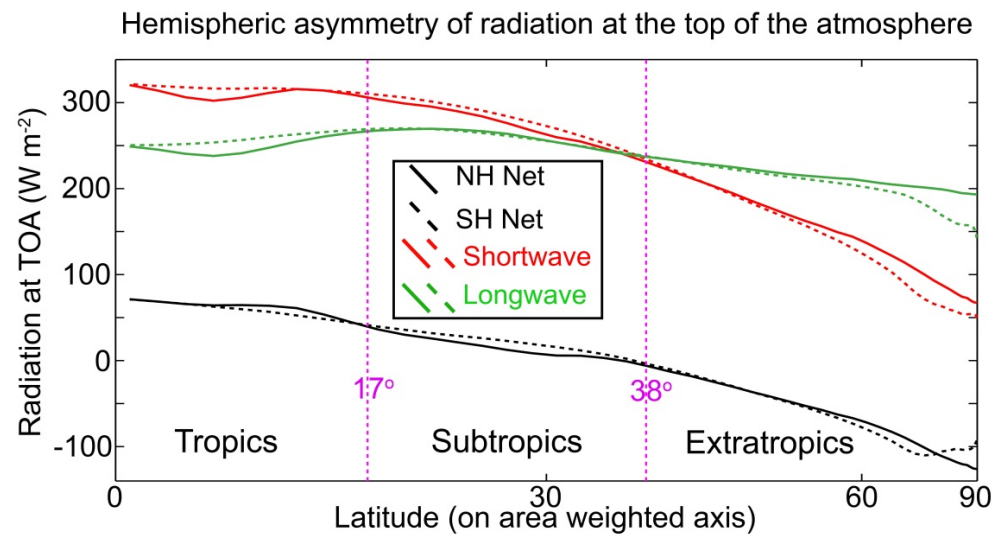
**B** Model Spread in ASR\* and Surface Albedo Contribution to ASR\*



**C** Model Spread in ASR\* and Surface Albedo Gradient

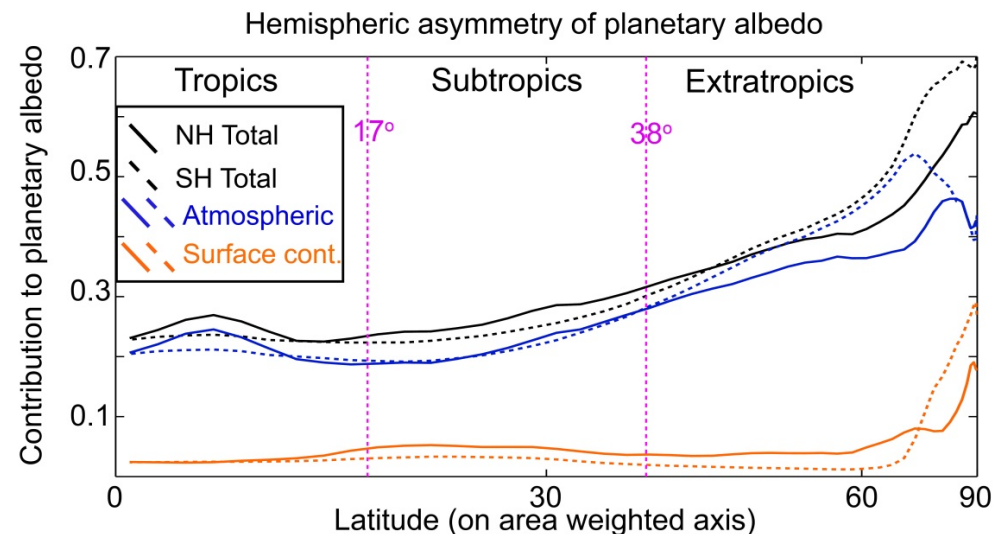


# Hemispheric Contrast Of TOA Radiation



NH reflects more SW radiation in the subtropical deserts. SH reflects more SW in the extratropics due to clouds in the Southern Ocean

The NH is warmer (more OLR), especially in the polar latitudes

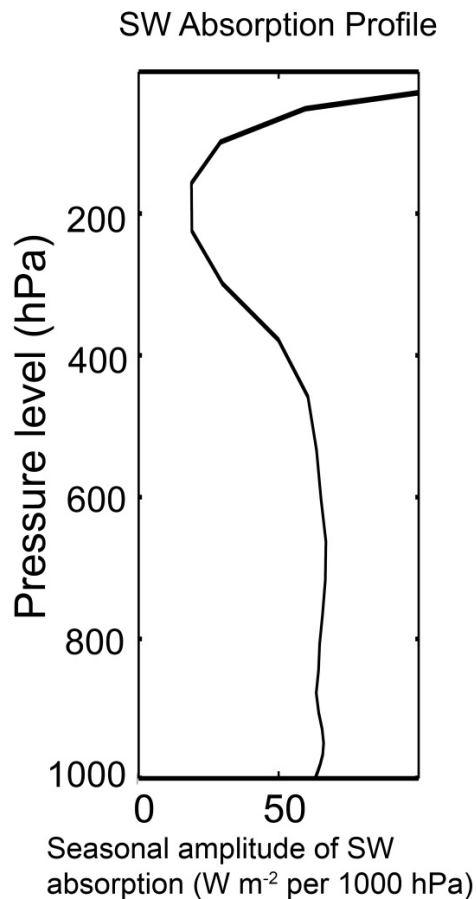


Planetary albedo is partitioned into cloud and surface contributions via the method of Donohoe and Battisti (2011)



# Vertical structure of SW absorption

GFDL Model  
Seasonal amplitude  
In the extratropics



Chou Lee (1996)  
Water vapor absorption  
In the summer

